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**UNIQUENESS OF GROUND STATES
FOR PSEUDORELATIVISTIC HARTREE EQUATIONS**

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UNIQUENESS OF GROUND STATES FOR PSEUDORELATIVISTIC HARTREE EQUATIONS

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We prove uniqueness of ground states $Q \in H^{1/2}(\mathbb{R}^3)$ for the pseudorelativistic Hartree equation,

$$\sqrt{-\Delta + m^2} Q - (|x|^{-1} * |Q|^2) Q = -\mu Q,$$

in the regime of Q with sufficiently small L^2 -mass. This result shows that a uniqueness conjecture by Lieb and Yau [1987] holds true at least for $N = \int |Q|^2 \ll 1$ except for at most countably many N .

Our proof combines variational arguments with a nonrelativistic limit, leading to a certain Hartree-type equation (also known as the Choquard–Pekard or Schrödinger–Newton equation). Uniqueness of ground states for this limiting Hartree equation is well-known. Here, as a key ingredient, we prove the so-called nondegeneracy of its linearization. This nondegeneracy result is also of independent interest, for it proves a key spectral assumption in a series of papers on effective solitary wave motion and classical limits for nonrelativistic Hartree equations.

1. Introduction

The pseudorelativistic Hartree energy functional, given (in appropriate units) by

$$\mathcal{E}(\psi) = \int_{\mathbb{R}^3} \bar{\psi} \sqrt{-\Delta + m^2} \psi - \frac{1}{2} \int_{\mathbb{R}^3} (|x|^{-1} * |\psi|^2) |\psi|^2, \quad (1-1)$$

arises in the mean-field limit of a quantum system describing many self-gravitating, relativistic bosons with rest mass $m > 0$. Such a physical system is often referred to as a *boson star*, and various models for these — at least theoretical — objects have attracted a great deal of attention in theoretical and numerical astrophysics over the past years.

In order to gain some rigorous insight into the theory of boson stars, it is of particular interest to study ground states (that is, minimizers) for the variational problem

$$E(N) = \inf \left\{ \mathcal{E}(\psi) : \psi \in H^{1/2}(\mathbb{R}^3) \text{ and } \int_{\mathbb{R}^3} |\psi|^2 = N \right\}, \quad (1-2)$$

where the parameter $N > 0$ plays the role of the stellar mass. Provided that problem (1-2) has indeed a ground state $Q \in H^{1/2}(\mathbb{R}^3)$, one readily finds that it satisfies the *pseudorelativistic Hartree equation*,

$$\sqrt{-\Delta + m^2} Q - (|x|^{-1} * |Q|^2) Q = -\mu Q, \quad (1-3)$$

with $\mu = \mu(Q) \in \mathbb{R}$ being some Lagrange multiplier.

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In fact, the existence of symmetric-decreasing ground states $Q = Q^*(|x|) \geq 0$ minimizing (1-2) was first proven by Lieb and Yau [1987], where the authors also conjectured that uniqueness holds true in the following sense. For each $N > 0$, the variational problem (1-2) has at most one symmetric-decreasing ground state. If true, this result further implies, by strict rearrangement inequalities, that we have indeed uniqueness of all the ground states of (1-2) for each $N > 0$, up to phase and translation.

However, the nonlocality of $\sqrt{-\Delta + m^2}$ as well as the convolution-type nonlinearity both complicate the analysis of the pseudorelativistic Hartree equation (1-3) in a substantial way. In particular, the set of its radial solutions is not amenable to ODE techniques (for example, shooting arguments and comparison principles) which are key arguments for proving uniqueness of ground states for nonlinear Schrödinger equations (NLS) with local nonlinearities; see [Peletier and Serrin 1983; McLeod and Serrin 1987; Kwong 1989; McLeod 1993].

A further complication in the analysis of (1-3) stems from the fact that there are no simple scaling arguments that relate ground states with different N , due to the presence of $m > 0$. Indeed, this lack of a simple scaling mechanism is essential for the existence of a critical stellar mass $N_* > 0$; see Theorem 1.

As a first step towards proving uniqueness of ground states for (1-2), we present Theorem 2 below, which shows that ground states for problem are indeed unique (modulo translation and phase) for all sufficiently small $N > 0$ except for at most countably many. Our proof uses variational arguments combined with a nonrelativistic limit, leading to the nonlinear Hartree equation (also called Choquard–Pekar or Schrödinger–Newton equation) given by

$$-\frac{1}{2m}\Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2)Q_\infty = -\lambda Q_\infty. \quad (1-4)$$

It is known this equation has a unique radial, positive solution $Q_\infty \in H^1(\mathbb{R}^3)$ for $\lambda > 0$ given; see [Lieb 1977] and Appendix A.

In the present paper, we prove (as a key ingredient) that $Q_\infty \in H^1(\mathbb{R}^3)$ has a *nondegenerate linearization*. By this we mean that the linearization of (1-4) around Q_∞ has a nullspace that is entirely due to the equation’s invariance under phase and translation transformation; see Theorem 4 below and its remarks for a precise statement. In particular, we show that the linear operator L_+ given by

$$L_+\xi = -\frac{1}{2m}\Delta\xi + \lambda\xi - (|x|^{-1} * |Q_\infty|^2)\xi - 2Q_\infty(|x|^{-1} * (Q_\infty\xi)) \quad (1-5)$$

satisfies

$$\ker L_+ = \text{span} \{\partial_{x_1} Q_\infty, \partial_{x_2} Q_\infty, \partial_{x_3} Q_\infty\}. \quad (1-6)$$

Furthermore, by a perturbation argument, we conclude an analogous nondegeneracy result for ground states of the pseudorelativistic Hartree equation (1-3) with sufficiently small L^2 -mass; see Theorem 3 below.

In addition to being a mere technical key fact proven in this paper, the nondegeneracy result for (1-4) is also of independent interest. For example, it proves a key spectral assumption in a series of papers on effective solitary wave motion and classical limits for Hartree equations; see [Fröhlich et al. 2002; 2004; Jonsson et al. 2006; Abou Salem 2007] and also the remark following Theorem 4. Another very recent application of the nondegeneracy result (1-6) is presented in [Krieger et al. 2008], where two soliton solutions to the time-dependent version of (1-4) are constructed.

In the context of ground states for NLS with *local nonlinearities*, the nondegeneracy of linearizations is a well-known fact (see [Weinstein 1985; Chang et al. 2007]) and it plays a central role in the stability analysis of solitary waves for NLS. However, the arguments for NLS with local nonlinearities make use of Sturm–Liouville theory, which, by contrast, is not applicable to L_+ given by (1-5) due to its nonlocal character. For more details, we refer to Section 7 below.

Apart from their minimizing property, the ground states for (1-2) also play an important role for the *time-dependent pseudorelativistic Hartree equation*,

$$i \partial_t \psi = \sqrt{-\Delta + m^2} \psi - (|x|^{-1} * |\psi|^2) \psi, \quad (1-7)$$

with the wave field $\psi : [0, T) \times \mathbb{R}^3 \rightarrow \mathbb{C}$. Clearly, Equation (1-7) has solitary wave solutions

$$\psi(t, x) = e^{it\mu} Q(x), \quad (1-8)$$

whenever $Q \in H^{1/2}(\mathbb{R}^3)$ is a nontrivial solution to (1-3). Let us also mention that the dispersive nonlinear PDE (1-7) exhibits a rich variety of phenomena, such as stable and unstable traveling solitary waves, as well as finite-time blowup solutions indicating the “gravitational collapse” of a boson star; see [Fröhlich et al. 2007a; 2007b; Fröhlich and Lenzmann 2007]. For well-posedness results concerning (1-7) and its rigorous derivation from many-body quantum mechanics, we refer to [Cho and Ozawa 2006; Lenzmann 2007] and [Elgart and Schlein 2007], respectively.

For the reader’s convenience, we conclude our introduction by summarizing the existence result about ground states for problem (1-2) along with a list of their basic properties.

Theorem 1 (Existence and properties of ground states). *Suppose that $m > 0$ holds in (1-1). Then there exists a universal constant $N_* > 4/\pi$ (independent of m) such that the following holds.*

(i) (Existence) *There exists a ground state $Q \in H^{1/2}(\mathbb{R}^3)$ for problem (1-2) if and only if*

$$0 < N < N_*.$$

Moreover, the function Q satisfies the pseudorelativistic Hartree equation (1-3) in the sense of distributions with some Lagrange multiplier $\mu \in \mathbb{R}$.

- (ii) (Smoothness and exponential decay) *Any ground state Q belongs to $H^s(\mathbb{R}^d)$ for all $s \geq 0$ and $e^{+\delta|x|} Q \in L^\infty(\mathbb{R}^3)$ for some $\delta = \delta(Q) > 0$.*
- (iii) (Radiality and strict positivity) *Any ground state Q is equal to its spherical-symmetric rearrangement $Q^*(|x|)$ up to phase and translation. Moreover, we have $Q^*(|x|) > 0$ for all $x \in \mathbb{R}^3$.*

Remark. For the proofs of (i) and (ii)–(iii), we refer to [Lieb and Yau 1987] and [Lenzmann 2006; Fröhlich et al. 2007a], respectively. In physical terms, the constant $N_* > 0$ can be regarded as the “Chandrasekhar limit mass” of a pseudorelativistic boson star.

2. Main results

We now state our first main result concerning the uniqueness of ground states for the pseudorelativistic Hartree equation (1-3).

Theorem 2 (Uniqueness of ground states for $N \ll 1$). *Assume that $m > 0$ holds in (1-1). Then, for $0 < N \ll 1$, we have uniqueness of ground states for problem (1-2) up to phase and translation whenever $E'(N)$ exists. In particular, the symmetric-decreasing ground state $Q = Q^* \in H^{1/2}(\mathbb{R}^3)$ minimizing (1-2) is unique for such $N > 0$.*

Remarks. (1) Since it is known from [Lieb and Yau 1987] that the ground state energy $E(N)$ is strictly concave, the derivative $E'(N)$ exists for all $N \in (0, N_*)$, except on a subset Σ which is at most countable. In particular, it is easy to see that the Lagrange multiplier μ is unique for such $N \in (0, N_*) \setminus \Sigma$, in the sense that μ only depends on Q through $N = \int |Q|^2$. Our argument to prove Theorem 2 has to avoid the “exceptional” set Σ . A natural conjecture would be that $\Sigma = \emptyset$ holds.

(2) It would be desirable to extend this uniqueness result (whose proof partly relies on perturbative arguments) to the whole range $0 < N < N_*$ of existence; or, more interestingly, to disprove uniqueness for some $N > 0$ sufficiently large.

(3) By definition, ground states for the pseudorelativistic Hartree equation (1-2) are always minimizers for the variational problem (1-2). In principle, we cannot exclude the possibility that (1-3) has a positive solution without being a minimizer for (1-2).

(4) To the author’s knowledge, this is the first uniqueness result for ground states that solve a nonlinear pseudo-differential equation in space dimension $n > 1$. In fact, apart from a very special case arising in $n = 1$ dimensions for solitary waves solving Benjamin–Ono-type equations (see [Amick and Toland 1991; Albert 1995]), nothing seems to be known, for instance, about uniqueness of ground states $\varphi \in H^s(\mathbb{R}^n)$ for nonlinear equations involving the fractional Laplacian $(-\Delta)^{s/2}\varphi + f(\varphi) = -\mu\varphi$, where $f(\varphi)$ denotes some nonlinearity and $\mu \in \mathbb{R}$ is given. The author plans to pursue this question in future work.

(5) If $m = 0$ vanishes, we have existence of ground states for problem (1-2) if and only if $N = N_*$. In what follows, we shall exclusively deal with the physically relevant case where $m > 0$ holds. Nevertheless, it remains an interesting open question whether uniqueness of ground states also holds for $m = 0$, since the methods developed here are clearly not applicable to this limiting case.

Our next result proves a so-called nondegeneracy condition, which was introduced in [Fröhlich et al. 2007b] as a spectral assumption supported by numerical evidence. There, the effective motion of solitary waves for (1-7) with an slowly varying external potential was studied. Furthermore, the following nondegeneracy result allows us to give an unconditional proof for the *cylindrical symmetry of traveling solitary waves* for the time-dependent pseudorelativistic Hartree equation (1-7); see [Fröhlich et al. 2007b] for more details. The precise nondegeneracy statement reads as follows.

Theorem 3 (Nondegeneracy of ground states for $N \ll 1$). *Let $m > 0$ in (1-1) and suppose that $Q = Q^*$ is a symmetric-decreasing ground state for problem (1-2) with Lagrange multiplier $\mu \in \mathbb{R}$. Furthermore, we consider the linear operator L_+ given by*

$$L_+\zeta = (\sqrt{-\Delta + m^2} + \mu)\zeta - (|x|^{-1} * |Q|^2)\zeta - 2Q(|x|^{-1} * (Q\zeta)),$$

acting on $L^2(\mathbb{R}^3)$ with domain $H^1(\mathbb{R}^3)$. Then, for $0 < N \ll 1$, the operator L_+ is nondegenerate, that is, its kernel satisfies

$$\ker L_+ = \text{span} \{ \partial_{x_1} Q, \partial_{x_2} Q, \partial_{x_3} Q \}.$$

Remarks. (1) This completely characterizes the kernel of the linearization of the pseudorelativistic Hartree equation (1-3) around ground state $Q = Q^*$ with $\int |Q|^2 \ll 1$. Note that, due to the presence of $|Q|^2$ in the nonlinearity, the linearized operator is not \mathbb{C} -linear. See also the remark following [Theorem 4](#) below for more details on the analogous statement for the nonrelativistic equation (1-4).

(2) The nondegeneracy of L_+ holds for all $N = \int |Q|^2 \ll 1$. The extra condition that $E'(N)$ exists, which is present in [Theorem 2](#), is not needed here.

In order to prove [Theorem 3](#), we first have to show the nondegeneracy for the linearization around the ground state $Q_\infty \in H^1(\mathbb{R}^3)$ solving the nonrelativistic Hartree equation (1-4). As mentioned before, this spectral result is of independent interest, since it proves a key assumption in [[Fröhlich et al. 2002](#); [Fröhlich et al. 2004](#); [Jonsson et al. 2006](#); [Abou Salem 2007](#)]. See also [[Krieger et al. 2008](#)], where the following nondegeneracy result is needed. Hence we record this fact about (1-4) as one of our main results.

Theorem 4 (Nondegeneracy for Q_∞). *Let $m > 0$ and $\lambda > 0$ be given. Furthermore, suppose that $Q_\infty \in H^1(\mathbb{R}^3)$ is the unique radial, positive solution to the nonrelativistic Hartree equation (1-4). Then the linear operator L_+ given by*

$$L_+\zeta = -\frac{1}{2m}\Delta\zeta + \lambda\zeta - (|x|^{-1} * |Q_\infty|^2)\zeta - 2Q_\infty(|x|^{-1} * (Q_\infty\zeta)) \quad (2-1)$$

acting on $L^2(\mathbb{R}^3)$ with domain $H^2(\mathbb{R}^3)$, satisfies

$$\ker L_+ = \text{span} \{ \partial_{x_1} Q_\infty, \partial_{x_2} Q_\infty, \partial_{x_3} Q_\infty \}. \quad (2-2)$$

Remarks. (1) The linearized operator L for (1-4) at Q_∞ is found to be

$$Lh = -\frac{1}{2m}\Delta h + \lambda h - (|x|^{-1} * |Q_\infty|^2)h - Q_\infty(|x|^{-1} * (Q_\infty(h + \bar{h}))).$$

It is convenient to view the operator L (which is not \mathbb{C} -linear) as acting on

$$\begin{pmatrix} \text{Re } h \\ \text{Im } h \end{pmatrix},$$

so that it can be written as

$$L = \begin{pmatrix} L_+ & 0 \\ 0 & L_- \end{pmatrix}.$$

Here L_+ is as in [Theorem 4](#) above, and L_- is the (local) operator

$$L_- = -\frac{1}{2m}\Delta + \lambda - (|x|^{-1} * |Q_\infty|^2).$$

It is easy to see that $\ker L_- = \text{span} \{Q_\infty\}$ holds. Hence, by [Theorem 4](#), we obtain

$$\ker L = \text{span} \left\{ \begin{pmatrix} \partial_{x_1} Q_\infty \\ 0 \end{pmatrix}, \begin{pmatrix} \partial_{x_2} Q_\infty \\ 0 \end{pmatrix}, \begin{pmatrix} \partial_{x_3} Q_\infty \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ Q_\infty \end{pmatrix} \right\}.$$

(2) The precise knowledge of $\ker L$ implies, by well-known arguments along the lines for NLS with local nonlinearities (given in [[Weinstein 1985](#)]), the following coercivity estimate: There is a constant

$\delta > 0$ such that

$$\langle f, L_+ f \rangle + \langle g, L_- g \rangle \geq \delta (\|f\|_{H^1}^2 + \|g\|_{H^1}^2),$$

when $f \perp \text{span}\{Q_\infty, x_i Q_\infty\}_{i=1}^3$ and $g \perp \text{span}\{2Q_\infty + r\partial_r Q_\infty, \partial_{x_i} Q_\infty\}_{i=1}^3$, which means that (f, g) is *symplectically orthogonal* to the ‘‘soliton manifold’’ generated by Q_∞ ; see, for example, [Fröhlich et al. 2004]. This coercivity estimate plays a central role in the stability analysis of solitary waves for NLS-type equations and their effective motion in an external potential; see, for example, [Weinstein 1985; Bronski and Jerrard 2000; Fröhlich et al. 2004; 2007b; Jonsson et al. 2006; Abou Salem 2007; Holmer and Zworski 2008].

Organization of the paper. This paper is structured as follows. In Section 3, we study the nonrelativistic limit of ground states for a dimensionalized version of the variational problem (1-2). In Section 4, we prove a nondegeneracy result for the nonrelativistic ground state $Q_\infty \in H^1(\mathbb{R}^3)$ in the radial setting. Then, in Section 5, we establish a local uniqueness result around $Q_\infty \in H^1(\mathbb{R}^3)$ by means of an implicit-function-type argument.

We prove Theorem 2 in Section 6, and Theorems 3 and 4 in Section 7. Appendices A and B collect some auxiliary results and we also give a uniqueness proof for the ground state $Q_\infty \in H^1(\mathbb{R}^3)$, which differs from [Lieb 1977] in certain ways.

Notation and conventions. As usual $H^s(\mathbb{R}^n)$ stands for the inhomogeneous Sobolev space of order $s \in \mathbb{R}$, equipped with norm $\|f\|_{H^s} = \|\langle \nabla \rangle^s f\|_{L^2}$, where $\langle \nabla \rangle$ is defined via its multiplier $\langle \xi \rangle = (1 + \xi^2)^{1/2}$ in the Fourier domain. Also, we shall make use of the space of radial and real-valued functions that belong to $H^1(\mathbb{R}^3)$, which we denote by

$$H_r^1(\mathbb{R}^3) = \{f : f \in H^1(\mathbb{R}^3), f \text{ is radial and real-valued}\}.$$

With the usual abuse of notation we shall write both $f(x)$ and $f(r)$, with $r = |x|$, for radial functions f on \mathbb{R}^n . For any measurable function $f : \mathbb{R}^n \rightarrow \mathbb{C}$ that vanishes at infinity, we denote its symmetric-decreasing rearrangement by $f^* = f^*(r) \geq 0$.

Throughout this paper, we assume that the mass parameter $m > 0$ in (1-1) is strictly positive, which is the physically relevant case.

For the reader’s orientation, we mention that our definition of $\mathcal{E}(\psi)$ in (1-1) differs from the conventions in [Lieb and Yau 1987; Fröhlich et al. 2007a] by an inessential factor of 2 and by the fact that we use $\sqrt{-\Delta + m^2}$ instead of $\sqrt{-\Delta + m^2} - m$. Obviously, these slight alterations in our definition of $\mathcal{E}(\psi)$ do not affect any results on (1-2) that are derived or quoted in the present paper.

Finally, we point out that the function $Q_\infty \in H_r^1(\mathbb{R}^3)$, which denotes the unique ground state for (1-4), appears throughout the paper. However, for the sake of simple notation, we shall also denote all its rescaled copies $aQ_\infty(b \cdot)$, with $a > 0$ and $b > 0$, simply by Q_∞ , whenever there is no source of confusion.

3. Nonrelativistic limit

As a preliminary step towards the proof of Theorems 2 and 3, we study the nonrelativistic limit of ground states for the pseudorelativistic Hartree energy functional. More precisely, we reinstall the speed of light

$c > 0$ into $\mathcal{E}(\psi)$ defined in (1-1), which yields the c -depending Hartree energy functional

$$\mathcal{E}_c(\psi) = \int_{\mathbb{R}^3} \bar{\psi} \sqrt{-c^2 \Delta + m^2 c^4} \psi - \frac{1}{2} \int_{\mathbb{R}^3} (|x|^{-1} * |\psi|^2) |\psi|^2. \quad (3-1)$$

An elementary calculation shows that, for any $\psi \in H^{1/2}(\mathbb{R}^3)$,

$$\mathcal{E}(\psi) = c^{-3} \mathcal{E}_c(\tilde{\psi}), \quad \text{with } \psi(x) = c^{-2} \tilde{\psi}(c^{-1}x). \quad (3-2)$$

Thus we immediately find the following equivalence.

Lemma 1. *Let $c > 0$ and $N > 0$. Then $\tilde{Q} \in H^{1/2}(\mathbb{R}^3)$ minimizes $\mathcal{E}_c(\psi)$ subject to $\int |\psi|^2 = N$ if and only if $Q = c^{-2} \tilde{Q}(c^{-1} \cdot)$ minimizes $\mathcal{E}(\psi)$ subject to $\int |\psi|^2 = c^{-1} N$.*

In particular, we have existence of ground states for $\mathcal{E}_c(\psi)$ subject to $\int |\psi|^2 = N$ if and only if $0 < N < cN_$ holds, where $N_* > 4/\pi$ denotes the same universal constant as in [Theorem 1](#).*

We now study the behavior of ground states Q_c for $\mathcal{E}_c(\psi)$ as $c \rightarrow \infty$ with $\int_{\mathbb{R}^3} |Q_c|^2 = N$ being fixed. By [Lemma 1](#), this is equivalent (after a suitable rescaling) to studying ground states for $\mathcal{E}(\psi)$ with $\int |\psi|^2 = N$ as $N \rightarrow 0$. However, the following analysis turns out to be more transparent when working with $c > 0$ as a parameter and sending c to infinity. Concerning the nonrelativistic limit $c \rightarrow \infty$ of ground states for $\mathcal{E}_c(\psi)$, we have the following result.

Proposition 1. *Let $m > 0$ and $N > 0$ be given, and suppose that $c_n \rightarrow \infty$ as $n \rightarrow \infty$. Furthermore, we assume that $\{Q_{c_n}\}_{n=1}^\infty$ is a sequence of symmetric-decreasing ground states such that $\int_{\mathbb{R}^3} |Q_{c_n}|^2 = N$ for all $n \geq 1$, and each $Q_{c_n} \in H^{1/2}(\mathbb{R}^3)$ minimizes $\mathcal{E}_{c_n}(\psi)$ subject to $\int_{\mathbb{R}^3} |\psi|^2 = N$. Finally, let $\{\mu_{c_n}\}_{n=1}^\infty$ denote the sequence of Lagrange multipliers corresponding to $\{Q_{c_n}\}_{n=1}^\infty$.*

Then the following holds:

$$\begin{aligned} Q_{c_n} &\rightarrow Q_\infty \quad \text{in } H^1(\mathbb{R}^3) & \text{as } n \rightarrow \infty, \\ -\mu_{c_n} - mc_n^2 &\rightarrow -\lambda & \text{as } n \rightarrow \infty, \end{aligned}$$

where $Q_\infty \in H^1(\mathbb{R}^3)$ is the unique radial, positive solution to

$$-\frac{1}{2m} \Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2) Q_\infty = -\lambda Q_\infty, \quad (3-3)$$

such that $\int_{\mathbb{R}^3} |Q_\infty|^2 = N$. Here $\lambda > 0$ is determined through $Q_\infty = Q_\infty^* \in H^1(\mathbb{R}^3)$, which is the unique symmetric-decreasing minimizer of the variational problem

$$E_{\text{nr}}(N) = \inf \left\{ \mathcal{E}_{\text{nr}}(\psi) : \psi \in H^1(\mathbb{R}^3) \text{ and } \int_{\mathbb{R}^3} |\psi|^2 = N \right\}, \quad (3-4)$$

where

$$\mathcal{E}_{\text{nr}}(\psi) = \frac{1}{2m} \int_{\mathbb{R}^3} |\nabla \psi|^2 - \frac{1}{2} \int_{\mathbb{R}^3} (|x|^{-1} * |\psi|^2) |\psi|^2. \quad (3-5)$$

Remarks. (1) A similar result for the nonrelativistic limit of ground states (and excited states) solving the Dirac–Fock equations can be found in [\[Esteban and Séré 2001\]](#). However, unlike the Dirac–Fock and Hartree–Fock energy functionals in atomic physics treated in [\[Esteban and Séré 2001\]](#), the energy functional in (3-1) is not weakly lower semicontinuous due to its attractive potential term. Therefore, an a priori bound on the sequence of Lagrange multipliers μ_{c_n} (away from the essential spectrum of the

limiting equation) is not sufficient to conclude strong convergence. To deal with this, we also have to use the radial symmetry of the Q_{c_n} in order to prove strong convergence.

(2) The uniqueness of the symmetric-decreasing ground state for problem (3-4) was proven by Lieb [1977]. For the reader's convenience, we provide a (partly different) proof of this fact in Appendix A.

3.1. Proof of Proposition 1. We begin with some auxiliary results.

Lemma 2. *Let $\{\mu_{c_n}\}_{n=1}^{\infty}$ be as in Proposition 1. Then there exist constants $\delta_1 > 0$ and $\delta_2 > 0$ such that*

$$mc_n^2 - \delta_1 \leq -\mu_{c_n} \leq mc_n^2 - \delta_2, \quad \text{for all } n \geq n_0,$$

where $n_0 \gg 1$ is some number.

Proof. The existence of $\delta_2 > 0$ can be deduced as follows. The Euler–Lagrange equation for Q_{c_n} reads

$$\sqrt{-c_n^2 \Delta + m^2 c_n^4} Q_{c_n} - (|x|^{-1} * |Q_{c_n}|^2) Q_{c_n} = -\mu_{c_n} Q_{c_n}, \quad (3-6)$$

which upon multiplication with Q_{c_n} and integration gives us

$$\mathcal{E}_{c_n}(Q_{c_n}) - \frac{1}{2} \int_{\mathbb{R}^3} (|x|^{-1} * |Q_{c_n}|^2) |Q_{c_n}|^2 = -\mu_{c_n} N. \quad (3-7)$$

Next, we recall the operator inequality

$$\sqrt{-c^2 \Delta + m^2 c^4} \leq -\frac{1}{2m} \Delta + mc^2,$$

which directly follows in the Fourier domain and the fact that $\sqrt{1+t} \leq t/2 + 1$ holds for all $t \geq 0$. Hence we have that $\mathcal{E}_{c_n}(Q_{c_n}) \leq \mathcal{E}_{\text{nr}}(Q_{c_n}) + Nm c_n^2$. Furthermore, since Q_{c_n} is a ground state for $\mathcal{E}_{c_n}(\psi)$, we deduce

$$\mathcal{E}_{c_n}(Q_{c_n}) \leq E_{\text{nr}}(N) + Nm c_n^2,$$

with $E_{\text{nr}}(N)$ defined in (3-4), so that (3-7) gives us

$$-\mu_{c_n} N \leq E_{\text{nr}}(N) + Nm c_n^2.$$

From [Lieb 1977] we know that $E_{\text{nr}}(N) < 0$ and thus $\delta_2 = -E_{\text{nr}}(N)/N > 0$ is a legitimate choice.

To prove the existence of $\delta_1 > 0$, we observe that each $Q_{c_n} \geq 0$ is the ground state of the “relativistic” Schrödinger operator

$$H_{c_n} = \sqrt{-c_n^2 \Delta + m^2 c_n^4} - (|x|^{-1} * |Q_{c_n}|^2).$$

Since all Q_{c_n} are radial functions with $\|Q_{c_n}\|_{L^2}^2 = N$ for all $n \geq 1$, we can invoke Newton's theorem to find

$$\int_{\mathbb{R}^3} \frac{|Q_{c_n}(y)|^2}{|x-y|} dy \leq \frac{N}{|x|}.$$

By the min-max principle, we infer the lower bound

$$-\mu_{c_n} \geq \inf \sigma(\bar{H}_{c_n})$$

where

$$\bar{H}_{c_n} = \sqrt{-c_n^2 \Delta + m^2 c_n^4} - \frac{N}{|x|}.$$

From [Herbst 1977] and reinstalling the speed of light $c > 0$ there, we recall that we have $\inf \sigma(\overline{H}_{c_n}) > -\infty$ if and only if $N < (2/\pi)c_n$. Thus \overline{H}_{c_n} is bounded below for $n \gg 1$ and, moreover, we have an explicit lower bound (see [Herbst 1977] again) given by

$$\inf \sigma(\overline{H}_{c_n}) \geq mc_n^2 \sqrt{1 - \left(\frac{\pi N}{2c_n}\right)^2}.$$

Since $\sqrt{1-x^2} \geq 1-x^2$ for $|x| \leq 1$, we conclude

$$-\mu_{c_n} \geq mc_n^2 \left(1 - \left(\frac{\pi N}{2c_n}\right)^2\right) = mc_n^2 - \frac{1}{4}m\pi^2 N^2, \quad \text{for all } n \geq n_0,$$

provided that $n_0 \gg 1$. By choosing $\delta_1 = \frac{1}{4}m\pi^2 N^2 > 0$, we complete the proof of Lemma 2. \square

Next, we derive an a priori bound on the sequence of ground states.

Lemma 3. *Let $\{Q_{c_n}\}_{n=1}^\infty$ be as in Proposition 1. Then there exists a constant $M > 0$ such that*

$$\|Q_{c_n}\|_{H^1} \leq M, \quad \text{for all } n \geq 1.$$

Proof. Since $\|Q_{c_n}\|_{L^2}^2 = N$ for all $n \geq 1$, we only have to derive a uniform bound for $\|\nabla Q_{c_n}\|_{L^2}$ which can be done as follows. From (3-6) we obtain

$$\begin{aligned} & c_n^2 \|\nabla Q_{c_n}\|_{L^2}^2 + m^2 c_n^4 \|Q_{c_n}\|_{L^2}^2 \\ &= \langle \sqrt{-c_n^2 \Delta + m^2 c_n^4} Q_{c_n}, \sqrt{-c_n^2 \Delta + m^2 c_n^4} Q_{c_n} \rangle \\ &\leq \mu_{c_n}^2 \langle Q_{c_n}, Q_{c_n} \rangle + 2|\mu_{c_n}| \langle Q_{c_n}, (|x|^{-1} * |Q_{c_n}|^2) Q_{c_n} \rangle + \langle Q_{c_n}, (|x|^{-1} * |Q_{c_n}|^2), (|x|^{-1} * |Q_{c_n}|^2) Q_{c_n} \rangle. \end{aligned}$$

To bound the terms on the right, we notice that Kato's inequality $|x|^{-1} \lesssim |\nabla|$ implies

$$\||x|^{-1} * |Q_{c_n}|^2\|_{L^\infty} \lesssim \langle Q_{c_n}, |\nabla| Q_{c_n} \rangle \lesssim \|Q_{c_n}\|_{L^2} \|\nabla Q_{c_n}\|_{L^2}.$$

Using this bound, Hölder's inequality, and the bound $|\mu_{c_n}| \leq mc_n^2$ for $n \gg 1$ from Lemma 2, we obtain

$$c_n^2 \|\nabla Q_{c_n}\|_{L^2}^2 \lesssim mc_n^2 N^{3/2} \|\nabla Q_{c_n}\|_{L^2} + N^2 \|\nabla Q_{c_n}\|_{L^2}^2,$$

for $n \gg 1$. Since $c_n \rightarrow \infty$ and N is fixed, we conclude that there exists $M > 0$ such that

$$\|\nabla Q_{c_n}\|_{L^2} \leq M$$

for $n \gg 1$. By choosing $M > 0$ possibly larger, we extend this bound to all $n \geq 1$. \square

We now come the proof of Proposition 1 itself. By the a priori bound in Lemma 3, we have (after possibly passing to a subsequence) that

$$Q_{c_n} \rightharpoonup Q_\infty \text{ in } H^1(\mathbb{R}^3) \text{ and } Q_{c_n}(x) \rightarrow Q_\infty(x) \text{ for a.e. } x \in \mathbb{R}^3 \text{ as } n \rightarrow \infty,$$

for some $Q_\infty \in H^1(\mathbb{R}^3)$. By radially and strict positivity of all the Q_{c_n} , it follows that $Q_\infty(|x|) \geq 0$ is a radial and nonnegative function. Furthermore, since $\{Q_{c_n}\}_{n=1}^\infty$ forms a sequence of radial functions on \mathbb{R}^3 with a uniform H^1 -bound, a classical result (see [Strauss 1977]) yields that

$$Q_{c_n} \rightarrow Q_\infty \text{ in } L^p(\mathbb{R}^3) \text{ as } n \rightarrow \infty \text{ for any } 2 < p < 6. \quad (3-8)$$

By [Lemma 2](#), we have that $\{-\mu_{c_n} - mc_n^2\}_{n=1}^\infty$ is a bounded sequence, which is also uniformly bounded away from 0. Hence extracting a suitable subsequence yields

$$\lim_{n \rightarrow \infty} (-\mu_{c_n} - mc_n^2) = -\lambda < 0, \quad (3-9)$$

for some $\lambda > 0$.

Using that $Q_{c_n} \rightharpoonup Q_\infty$ in H^1 and the strong convergence [\(3-8\)](#), we can pass to the limit in [\(3-6\)](#) and find that the radial, nonnegative function $Q_\infty \in H^1(\mathbb{R}^3)$ satisfies

$$-\frac{1}{2m} \Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2) Q_\infty = -\lambda Q_\infty \quad \text{in } H^{-1}(\mathbb{R}^3). \quad (3-10)$$

When taking this limit, we use the fact that

$$\lim_{n \rightarrow \infty} \left\langle f, \left(\sqrt{-c_n^2 \Delta + m^2 c_n^4} - mc_n^2 + \frac{1}{2m} \Delta \right) Q_{c_n} \right\rangle = 0 \quad \text{for all } f \in H^1(\mathbb{R}^3),$$

which is easy to verify for test functions $f \in C_0^\infty(\mathbb{R}^3)$ by taking the Fourier transform and using that

$$\sqrt{c_n^2 \xi^2 + m^2 c_n^4} - mc_n^2 - \frac{\xi^2}{2m} \rightarrow 0 \quad \text{for every } \xi \in \mathbb{R}^3 \text{ as } c_n \rightarrow \infty.$$

The claim above extends to all $f \in H^1(\mathbb{R}^3)$ by a simple density argument.

Next we prove that in fact $\int |Q_\infty|^2 = N$ holds, which a-posteriori would show that $Q_{c_n} \rightarrow Q_\infty$ strongly in $L^2(\mathbb{R}^3)$. To prove this claim, we note that [Equation \(3-6\)](#) and its limit [\(3-10\)](#) give us

$$(-\mu_{c_n} - mc_n^2)N = \frac{1}{2m} \int_{\mathbb{R}^3} |\nabla Q_{c_n}|^2 - \int_{\mathbb{R}^3} (|x|^{-1} * |Q_{c_n}|^2) |Q_{c_n}|^2 + r_n, \quad (3-11)$$

with $r_n \rightarrow 0$ as $n \rightarrow \infty$. Note that the right-hand side is not weakly lower semicontinuous (with respect to weak H^1 -convergence), unlike the case of atomic Hartree and Hartree–Fock energy functionals. To deal with the non weakly lower semicontinuous part given by the potential energy term, we use [\(3-8\)](#) again and the Hardy–Littlewood–Sobolev inequality. Then, by the weak lower semicontinuity of the kinetic energy term in [\(3-11\)](#), we deduce from [\(3-11\)](#) and [\(3-10\)](#) that

$$-\lambda N \geq \frac{1}{2m} \int_{\mathbb{R}^3} |\nabla Q_\infty|^2 - \int_{\mathbb{R}^3} (|x|^{-1} * |Q_\infty|^2) |Q_\infty|^2 = -\lambda \int_{\mathbb{R}^3} |Q_\infty|^2.$$

Because of $\lambda > 0$, we see that $\int |Q_\infty|^2 \geq N$ must hold. On the other hand, we have $N \geq \int |Q_\infty|^2$ by the weak L^2 -convergence. Thus we have $\int |Q_\infty|^2 = N$ and, consequently,

$$Q_{c_n} \rightarrow Q_\infty \text{ in } L^2(\mathbb{R}^3) \text{ as } n \rightarrow \infty. \quad (3-12)$$

By [Lemma 9](#) and a simple scaling argument, we see that Q_∞ is the unique radial, nonnegative solution to [\(3-10\)](#) with $\int |Q_\infty|^2 = N$. Here $\lambda > 0$ is determined through Q_∞ , and Q_∞ is in fact strictly positive.

It remains to show that

$$Q_{c_n} \rightarrow Q_\infty \text{ in } H^1(\mathbb{R}^3) \text{ as } n \rightarrow \infty. \quad (3-13)$$

To see this, we verify that $\{Q_{c_n}\}_{n=1}^\infty$ with $\int |Q_{c_n}|^2 = N$ furnishes a minimizing sequence for the nonrelativistic Hartree energy $\mathcal{E}_{\text{nr}}(\psi)$ subject to $\int |\psi|^2 = N$, that is, for problem [\(3-4\)](#). Indeed, using [\(3-11\)](#)

and (3-9) as well as the strong convergence (3-8) to pass to the limit in the potential energy, we deduce that

$$\mathcal{E}_{\text{nr}}(Q_{c_n}) \rightarrow -\lambda N + \frac{1}{2} \int_{\mathbb{R}^3} (|x|^{-1} * |Q_\infty|^2) |Q_\infty|^2 \quad \text{as } n \rightarrow \infty.$$

However, this limit for $\mathcal{E}_{\text{nr}}(Q_{c_n})$ is equal to $\mathcal{E}_{\text{nr}}(Q_\infty)$, as can be seen by multiplying (3-10) with Q_∞ and integrating. Hence $\{Q_{c_n}\}_{n=1}^\infty$ is a minimizing sequence for problem (3-4). Next, we notice that standard concentration-compactness methods yield relative compactness in $H^1(\mathbb{R}^3)$ for any radial minimizing sequence for problem (3-4), which has a unique radial, nonnegative minimizer Q_∞ . Therefore (after possibly passing to another subsequence) we deduce that (3-13) holds.

To conclude the proof of Proposition 1, we note that we have convergence along every subsequence because of the uniqueness of the limit point $Q_\infty \in H^1(\mathbb{R}^3)$. \square

4. Radial nondegeneracy of nonrelativistic ground states

We consider the linear operator

$$L_+ \zeta = -\frac{1}{2m} \Delta \zeta + \lambda \zeta - (|x|^{-1} * |Q_\infty|^2) \zeta - 2Q_\infty (|x|^{-1} * (Q_\infty \zeta)), \quad (4-1)$$

where $Q_\infty \in H^1(\mathbb{R}^3)$ is the radial, positive solution taken from Proposition 1. By standard arguments, it follows that L_+ is a self-adjoint operator acting on $L^2(\mathbb{R}^3)$ with domain $H^2(\mathbb{R}^3)$. In this section, we study the restriction of L_+ acting on $L^2_{\text{rad}}(\mathbb{R}^3)$ (that is, the radial L^2 -functions on \mathbb{R}^3).

As a main result, we prove the so-called nondegeneracy of L_+ on $L^2_{\text{rad}}(\mathbb{R}^3)$; that is, the triviality of its kernel.

Proposition 2. *For the linear operator L_+ be given by (4-1), we have*

$$\ker L_+ = \{0\} \quad \text{when } L_+ \text{ is restricted to } L^2_{\text{rad}}(\mathbb{R}^3).$$

Remark. (1) As shown in Section 7 below, we will see that the triviality of the kernel of L_+ on $L^2_{\text{rad}}(\mathbb{R}^3)$ implies

$$\ker L_+ = \text{span} \{ \partial_{x_1} Q_\infty, \partial_{x_2} Q_\infty, \partial_{x_3} Q_\infty \}. \quad (4-2)$$

For linearized operators arising from ground states for NLS with *local nonlinearities*, this fact is well-known; see [Chang et al. 2007; Weinstein 1985]. However, the proof given there *cannot be adapted* to L_+ given by (4-1) due to its nonlocal component. We refer to Section 7 for further details.

(2) Numerical evidence indicating that 0 is not an eigenvalue of L_+ when restricted to radial functions can be found in [Harrison et al. 2003].

4.1. Proof of Proposition 2. Suppose that $Q_\infty \in H^1(\mathbb{R}^3)$ is the unique radial, positive solution to (3-3) with $\int |Q_\infty|^2 = N$ for some $N > 0$ given. In what follows, it will be convenient and without loss of generality to assume that Q_∞ satisfies

$$-\Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2) Q_\infty = -Q_\infty, \quad (4-3)$$

which amounts to rescaling $Q_\infty(x) \mapsto aQ_\infty(bx)$ with suitable $a > 0$ and $b > 0$. Likewise, the linear operator L_+ then reads

$$L_+\zeta = -\Delta\zeta + \zeta - (|x|^{-1} * |Q_\infty|^2)\zeta - 2Q_\infty(|x|^{-1} * (Q_\infty\zeta)). \quad (4-4)$$

Recall that we restrict ourselves to radial $\zeta \in L^2_{\text{rad}}(\mathbb{R}^3)$. Therefore, we can rewrite the nonlocal term in L_+ by invoking Newton's theorem in \mathbb{R}^3 (see [Lieb and Loss 2001, Theorem 9.7]): For any radial function $\rho = \rho(|x|)$ such that $\rho \in L^1(\mathbb{R}^3, (1 + |x|)^{-1}dx)$, we have

$$-(|x|^{-1} * \rho)(r) = \int_0^r K(r, s)\rho(s) ds - \int_{\mathbb{R}^3} \frac{\rho(|x|)}{|x|}, \quad (4-5)$$

for $r = |x| \geq 0$, where $K(r, s)$ is given by

$$K(r, s) = 4\pi s \left(1 - \frac{s}{r}\right) \geq 0, \quad \text{for } r \geq s. \quad (4-6)$$

Since the ground state Q_∞ is exponentially decaying, we can apply Newton's theorem to $\rho = Q_\infty\zeta$ for any $\zeta \in L^2_{\text{rad}}(\mathbb{R}^3)$ and obtain the following result.

Lemma 4. *For any $\zeta \in L^2_{\text{rad}}(\mathbb{R}^3)$, we have*

$$L_+\zeta = \mathcal{L}_+\zeta - 2Q_\infty \left(\int_{\mathbb{R}^3} \frac{Q_\infty\zeta}{|x|} \right), \quad (4-7)$$

where \mathcal{L}_+ is given by

$$\mathcal{L}_+\zeta = -\Delta\zeta + \zeta - (|x|^{-1} * |Q_\infty|^2)\zeta + W\zeta, \quad (4-8)$$

with

$$(W\zeta)(r) = 2Q_\infty(r) \int_0^r K(r, s)Q_\infty(s)\zeta(s) ds. \quad (4-9)$$

The following auxiliary result shows exponential growth of solutions v to the linear equation $\mathcal{L}_+v = 0$.

Lemma 5. *Suppose the radial function $v = v(r)$ solves $\mathcal{L}_+v = 0$ with $v(0) \neq 0$ and $v'(0) = 0$. Then the function $v(r)$ has no sign change and $v(r)$ grows exponentially as $r \rightarrow \infty$. More precisely, for any $0 < \delta < 1$, there exist constants $C > 0$ and $R > 0$ such that*

$$|v(r)| \geq Ce^{+\delta r}, \quad \text{for all } r \geq R.$$

In particular, we have that $v \notin L^2_{\text{rad}}(\mathbb{R}^3)$.

Proof. Since $\mathcal{L}_+v = 0$ is a linear equation, we can assume without loss of generality that $v(0) > 0$; and moreover it is convenient to assume that $v(0) > Q_\infty(0)$ holds. Next, we write $\mathcal{L}_+v = 0$ as

$$v''(r) + \frac{2}{r}v'(r) = V(r)v(r) + W(r), \quad (4-10)$$

with

$$V(r) = 1 - (|x|^{-1} * |Q_\infty|^2)(r), \quad (4-11)$$

$$W(r) = 2Q_\infty(r) \int_0^r K(r, s)Q_\infty(s)v(s) ds. \quad (4-12)$$

Note that $Q_\infty(r)$ satisfies (4-10) with $W(r)$ being removed, that is,

$$Q_\infty''(r) + \frac{2}{r}Q_\infty'(r) = V(r)Q_\infty(r). \quad (4-13)$$

We now compare $v(r)$ and $Q_\infty(r)$ as follows. An elementary calculation, using equations (4-10) and (4-13), leads to the ‘‘Wronskian-type’’ identity

$$(r^2(Q_\infty v' - Q_\infty' v))' = r^2 Q_\infty W, \quad (4-14)$$

which, by integration, gives us

$$r^2(Q_\infty v' - Q_\infty' v)(r) = \int_0^r s^2 Q_\infty(s) W(s) ds. \quad (4-15)$$

Hence, while keeping in mind that $Q_\infty(r) > 0$, we find

$$r^2 \left(\frac{v(r)}{Q_\infty(r)} \right)' = \frac{1}{Q_\infty(r)^2} \int_0^r s^2 Q_\infty(s) W(s) ds. \quad (4-16)$$

From this identity we now claim that

$$v(r) > Q_\infty(r), \quad \text{for all } r \geq 0. \quad (4-17)$$

To see this, recall that $v(0) > Q_\infty(0)$ and, by continuity, we have that $v(r) > Q_\infty(r)$ for $r > 0$ sufficiently small. Suppose now, on the contrary to (4-17), that there is a first intersection at some positive $r = r_*$, say, so that $v(r_*) = Q_\infty(r_*)$. It is easy to see that the left-hand side of (4-16) (or equivalently (4-15)) has to be ≤ 0 at $r = r_*$. On the other hand, since $v(r) > Q_\infty(r) > 0$ on $[0, r_*)$, we conclude that the integral on right-hand side of (4-16) at $r = r_*$ must be strictly positive. This contradiction shows that (4-17) must hold. In particular, the function $v(r)$ never changes its sign.

Next, we insert the estimate (4-17) back into (4-16), which yields

$$r^2 \left(\frac{v(r)}{Q_\infty(r)} \right)'(r) \geq \frac{2}{Q_\infty(r)^2} \int_0^r s^2 Q_\infty(s)^2 \int_0^s K(s, t) Q_\infty(t)^2 dt ds. \quad (4-18)$$

We notice that $Q_\infty(r) > 0$ is the unique ground state for the Schrödinger operator

$$H = -\Delta + \tilde{V}, \quad \text{with } \tilde{V} = -|x|^{-1} * |Q_\infty|^2. \quad (4-19)$$

Since $HQ_\infty = -Q_\infty$ and \tilde{V} is a continuous function with $\tilde{V} \rightarrow 0$ as $|x| \rightarrow \infty$, standard arguments show that, for any $\varepsilon > 0$, there exists a constant $A_\varepsilon > 0$ such that

$$Q_\infty(r) \leq A_\varepsilon e^{-(1-\varepsilon)r}, \quad \text{for all } r \geq 0. \quad (4-20)$$

Furthermore, since $Q_\infty(r) > 0$ is the ground state of H , we can obtain the following lower bound: For any $\varepsilon > 0$, there exists a constant $B_\varepsilon > 0$ such that

$$Q_\infty(r) \geq B_\varepsilon e^{-(1+\varepsilon)r}, \quad \text{for all } r \geq 0. \quad (4-21)$$

For this classical result on ground states for Schrödinger operators. See, for example, [Carmona and Simon 1981, Theorem 3.2] where a probabilistic proof is given.

Now let $0 < \varepsilon < 1$ be given. Inserting the bounds (4-20) and (4-21) into Equation (4-18), we obtain

$$r^2 \left(\frac{v(r)}{Q(r)} \right)' (r) \geq C e^{(2-2\varepsilon)r} \int_0^r s^2 e^{-(2+2\varepsilon)s} \int_0^s K(s, t) e^{-(2+2\varepsilon)t} dt ds, \quad (4-22)$$

with some constant $C = C_\varepsilon > 0$ (we drop its dependence on ε henceforth). Since the double integral on the right-hand side converges as $r \rightarrow \infty$ to some finite positive value, there exists some $a > 0$ such that

$$r^2 \left(\frac{v(r)}{Q(r)} \right)' (r) \geq C e^{(2-2\varepsilon)r}, \quad \text{for all } r \geq a, \quad (4-23)$$

with some constant $C > 0$. Integrating this lower bound and using (4-21) again, we find that

$$v(r) \geq C \frac{e^{(1-3\varepsilon)r}}{r^2}, \quad \text{for all } r \geq R, \quad (4-24)$$

with some constants $C > 0$ and $R \gg 1$. Thus, for any $0 < \delta < 1$, we arrive at the claim of Lemma 5 by taking $0 < \varepsilon < \frac{1}{3}(1 - \delta)$ and choosing $C > 0$ appropriately. \square

With the help of Lemma 5 we are now able to prove the triviality of the kernel of L_+ in the radial sector.

Lemma 6. *For L_+ be given by (4-1), we have that $L_+\zeta = 0$ with $\zeta \in L_{\text{rad}}^2(\mathbb{R}^3)$ implies that $\zeta \equiv 0$.*

Proof. Suppose there exists $\zeta \in L_{\text{rad}}^2(\mathbb{R}^3)$ with $\zeta \not\equiv 0$ such that $L_+\zeta = 0$. Then, by Lemma 4, the function ζ solves the inhomogeneous problem

$$\mathcal{L}_+\zeta = 2\sigma Q_\infty, \quad \text{with } \sigma = \int_{\mathbb{R}^3} \frac{Q_\infty \zeta}{|x|}. \quad (4-25)$$

Therefore,

$$\zeta = v + w, \quad (4-26)$$

where w is any particular solution to (4-25) and v is some function such that $\mathcal{L}_+v = 0$. As shown below, it suffices to restrict ourselves to smooth v and w .

We shall now construct a smooth $w \in L_{\text{rad}}^2(\mathbb{R}^3)$ as follows. We define the smooth radial function

$$R = 2Q_\infty + r\partial_r Q_\infty \in L_{\text{rad}}^2(\mathbb{R}^3), \quad (4-27)$$

where a calculation shows that

$$L_+R = -2Q_\infty. \quad (4-28)$$

Furthermore, by applying Lemma 4 to R , we find

$$\mathcal{L}_+R = 2(\tau - 1)Q_\infty, \quad \text{with } \tau = \int_{\mathbb{R}^3} \frac{Q_\infty R}{|x|}. \quad (4-29)$$

Note that $\tau \neq 1$ must hold, for otherwise Lemma 5 with $v = R$ (and $v(0) = R(0) = Q(0) > 0$ and $v'(0) = R'(0) = 0$) would yield that $R \notin L_{\text{rad}}^2(\mathbb{R}^3)$, which is a contradiction. Thus we have found a smooth particular solution to (4-25) given by

$$w = \frac{\sigma}{\tau - 1} R \in L_{\text{rad}}^2(\mathbb{R}^3). \quad (4-30)$$

Further, we notice that $\zeta \in L_{\text{rad}}^2(\mathbb{R}^3)$ with $L_+\zeta = 0$ is smooth by bootstrapping this equation. Therefore, by (4-26), we conclude that v has to be smooth as well. Suppose that $v \equiv 0$. Then we have $\zeta = w$ and $\sigma \neq 0$ (since otherwise $w = 0 \neq \zeta$). This, however, contradicts that $L_+\zeta = 0$ and $L_+w = -2(\sigma/(\tau-1))Q_\infty \neq 0$.

Thus we see that $v \not\equiv 0$ in (4-26), where $v'(0) = 0$ by smoothness of v . Suppose now that $v(0) \neq 0$. Then Lemma 5 yields that $v \notin L_{\text{rad}}^2(\mathbb{R}^3)$, which contradicts (4-26) together with the fact that ζ and w both belong to $L_{\text{rad}}^2(\mathbb{R}^3)$. Finally, suppose that $v(0) = 0$ holds. Then v solves the equation $\mathcal{L}_+v = 0$ with initial data $v(0) = 0$ and $v'(0) = 0$. However, by a standard fixed point argument, we see that the linear integro-differential equation $\mathcal{L}_+v = 0$ with given initial data $v(0) \in \mathbb{R}$ and $v'(0) = 0$ has a unique solution. So $v(0) = 0$ and $v'(0) = 0$ implies that $v \equiv 0$. Again, we arrive at a contradiction as above. \square

Clearly, Lemma 6 completes the proof of Proposition 2. \square

5. Local uniqueness around Q_∞

Recall that $H_r^1(\mathbb{R}^3)$ denotes space of radial and real-valued functions that belong to $H^1(\mathbb{R}^3)$. By using Proposition 2, we can now prove the following local uniqueness result for a small neighborhood around Q_∞ in $H_r^1(\mathbb{R}^3)$.

Proposition 3. *Let $m > 0$ and $N > 0$ be given. Furthermore, suppose that $Q_\infty \in H_r^1(\mathbb{R}^3)$ is the unique radial, positive solution to*

$$-\frac{1}{2m}\Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2)Q_\infty = -\lambda Q_\infty, \quad (5-1)$$

with $\int |Q_\infty|^2 = N$, where $\lambda > 0$ is determined through Q_∞ . Then there exist constants $c_0 \gg 1$, $\varepsilon > 0$, and $\delta > 0$ such that the following holds. For any (c, μ) with

$$c \geq c_0, \quad -\lambda - \varepsilon \leq -\mu - mc^2 \leq -\lambda + \varepsilon,$$

the equation

$$\sqrt{-c^2\Delta + m^2c^4}Q - (|x|^{-1} * |Q|^2)Q = -\mu Q \quad (5-2)$$

has a unique solution $Q \in H_r^1(\mathbb{R}^3)$, provided that $\|Q - Q_\infty\|_{H^1} \leq \delta$.

5.1. Proof of Proposition 3. For $\beta \geq 0$ and $z > 0$, we define the map

$$G(u, \beta, z) = u + \mathcal{R}(\beta, z)g(u), \quad (5-3)$$

where we set

$$g(u) = -(|x|^{-1} * |u|^2)u, \quad (5-4)$$

and, for $\beta \geq 0$ and $z > 0$, we define the family of resolvents

$$\mathcal{R}(\beta, z) = \begin{cases} (-(1/2m)\Delta + z)^{-1} & \text{if } \beta = 0, \\ (\sqrt{-\beta^{-2}\Delta + m^2\beta^{-4}} - m\beta^{-2} + z)^{-1} & \text{if } \beta > 0. \end{cases} \quad (5-5)$$

By an elementary calculation, we verify the following equivalences:

$$Q \in H_r^1(\mathbb{R}^3) \text{ solves (5-1) if and only if } G(Q, 0, \lambda) = 0; \quad (5-6)$$

$$Q \in H_r^1(\mathbb{R}^3) \text{ solves (5-2) if and only if } G(Q, c^{-1}, \mu + mc^2) = 0. \quad (5-7)$$

To prove [Proposition 3](#), we now construct an implicit function-type argument for the map

$$G : H_r^1(\mathbb{R}^3) \times [0, \beta_0] \times [\lambda - \varepsilon, \lambda + \varepsilon] \rightarrow H_r^1(\mathbb{R}^3), \quad (5-8)$$

where $\beta_0 > 0$ and $\varepsilon > 0$ are small constants. To see that indeed $G(u, \beta, z) \in H_r^1(\mathbb{R}^3)$ for $u \in H_r^1(\mathbb{R}^3)$, we notice that $\mathcal{R}(\beta, z) : H_r^1(\mathbb{R}^3) \rightarrow H_r^1(\mathbb{R}^3)$, as can be seen by using the Fourier transform. That $g(u)$ maps $H_r^1(\mathbb{R}^3)$ into itself follows readily from the Hardy–Littlewood–Sobolev inequality and Sobolev embeddings. Hence (5-8) is indeed well-defined.

Next, we show that the derivative

$$\partial_u G(u, \beta, z) = 1 + \mathcal{R}(\beta, z) \partial_u g(u) : H_r^1(\mathbb{R}^3) \rightarrow H_r^1(\mathbb{R}^3) \quad (5-9)$$

depends continuously on (u, β, z) . Here $\partial_u g(u)$ acting on $\xi \in H_r^1(\mathbb{R}^3)$ is found to be

$$\partial_u g(u) \xi = -(|x|^{-1} * |u|^2) \xi - 2u(|x|^{-1} * (u \xi)). \quad (5-10)$$

By using the Hardy–Littlewood–Sobolev inequality and Sobolev embeddings, we obtain that

$$\|(\partial_u g(u_1) - \partial_u g(u_2)) \xi\|_{H^1} \lesssim (\|u_1\|_{H^1} + \|u_2\|_{H^1}) \|u_1 - u_2\|_{H^1} \|\xi\|_{H^1}; \quad (5-11)$$

see, for example, [\[Lenzmam 2007\]](#) for similar estimates proving Lipschitz continuity of $g(u)$. Using this estimate, we find for $u_1, u_2, \xi \in H_r^1(\mathbb{R}^3)$, $\beta_1, \beta_2 \in [0, \beta_0]$, and $z_1, z_2 > 0$,

$$\begin{aligned} & \|(\partial_u G(u_1, \beta_1, z_1) - \partial_u G(u_2, \beta_2, z_2)) \xi\|_{H^1} \\ & \leq \|(\mathcal{R}(\beta_1, z_1) - \mathcal{R}(\beta_2, z_2)) \partial_u g(u_1) \xi\|_{H^1} + \|\mathcal{R}(\beta_2, z_2) (\partial_u g(u_1) - \partial_u g(u_2)) \xi\|_{H^1} \\ & \lesssim \|\mathcal{R}(\beta_1, z_1) - \mathcal{R}(\beta_2, z_2)\|_{L^2 \rightarrow L^2} \|u_1\|_{H^1}^2 \|\xi\|_{H^1} \\ & \quad + \|\mathcal{R}(\beta_2, z_2)\|_{L^2 \rightarrow L^2} (\|u_1\|_{H^1} + \|u_2\|_{H^1}) \|u_1 - u_2\|_{H^1} \|\xi\|_{H^1}, \end{aligned} \quad (5-12)$$

where we also use the fact that $\|\mathcal{R}(\beta, z)\|_{H^s \rightarrow H^s} = \|\mathcal{R}(\beta, z)\|_{L^2 \rightarrow L^2}$ for any $s \in \mathbb{R}$, since $\mathcal{R}(\beta, z)$ commutes with $\langle \nabla \rangle$. Moreover, by using the Fourier transform, one verifies

$$\|\mathcal{R}(\beta_1, z_1) - \mathcal{R}(\beta_2, z_2)\|_{L^2 \rightarrow L^2} \rightarrow 0 \quad \text{as} \quad (\beta_1, z_1) \rightarrow (\beta_2, z_2), \quad (5-13)$$

for any $\beta_1, \beta_2 \geq 0$ and $z_1, z_2 > 0$. (For later use, we record that (5-13) also holds for complex $z_1, z_2 \in \mathbb{C} \setminus [0, \infty)$.) Going back to (5-12), we thus find

$$\|\partial_u G(u_1, \beta_1, z_1) - \partial_u G(u_2, \beta_2, z_2)\|_{H^1 \rightarrow H^1} \rightarrow 0$$

as $\|u_1 - u_2\|_{H^1} \rightarrow 0$ and $(\beta_1, z_1) \rightarrow (\beta_2, z_2)$. Hence $\partial_u G(u, \beta, z)$ depends continuously on (u, β, z) .

By [Proposition 2](#) and its following remark, we have that the radial restriction of the linearized operator L_+ around Q_∞ has trivial kernel. This implies that the compact operator $(-(1/(2m))\Delta + \lambda)^{-1} \partial_u g(Q_\infty)$ does not have -1 in its spectrum. Hence the inverse operator

$$(\partial_u G(Q_\infty, 0, \lambda))^{-1} : H_r^1(\mathbb{R}^3) \rightarrow H_r^1(\mathbb{R}^3) \quad (5-14)$$

exists. By the continuity of $\partial_u G(u, \beta, z)$ shown above, an appropriate version of an implicit function theorem (see, for example, [\[Chang 2005\]](#)) implies that, for $\beta_0 > 0$ and $\varepsilon > 0$ sufficiently small, there

exists a unique solution $Q = Q(\beta, z) \in H_r^1(\mathbb{R}^3)$ such that

$$G(Q(\beta, z), \beta, z) = 0 \quad \text{for } \beta \in [0, \beta_0] \text{ and } z \in [\lambda - \varepsilon, \lambda + \varepsilon] \quad (5-15)$$

with

$$\|Q(\beta, z) - Q_\infty\|_{H^1} \leq \delta \quad \text{for some } \delta > 0. \quad (5-16)$$

Moreover, the map $(\beta, z) \mapsto Q(\beta, z) \in H_r^1(\mathbb{R}^3)$ is continuous.

By setting $c_0 = \beta_0^{-1}$ and recalling the equivalence (5-7), we complete the proof of [Proposition 3](#). \square

6. Proof of [Theorem 2](#)

First, we notice that it is sufficient to prove uniqueness of symmetric-decreasing ground states for the variational problem (1-2), thanks to [Theorem 1](#) (iii). Next, we make use of the rescaling correspondence formulated in [Lemma 1](#), which relates ground states for the dimensionalized and de-dimensionalized Hartree energy functionals $\mathcal{E}_c(\psi)$ and $\mathcal{E}(\psi)$ defined in (3-1) and (1-1), respectively.

In what follows, we fix $\int |Q_c|^2 = 1$ and we suppose that $Q_c = Q_c^* \in H^{1/2}(\mathbb{R}^3)$ is a symmetric-decreasing ground state for $\mathcal{E}_c(\psi)$ subject to $\int |\psi|^2 = 1$. Recall from [Lemma 1](#) that Q_c indeed exists for $c \geq c_0$ with c_0 being a sufficiently large constant. Let $\mu(Q_c)$ denote the Lagrange multiplier associated to Q_c for $c \geq c_0$. We now claim that μ only depends on c except for some countable set, that is, we have

$$\mu(Q_c) = \mu(c), \quad \text{for } c \in (c_0, \infty) \setminus \Xi, \quad (6-1)$$

where Ξ is some countable set. To prove (6-1), we argue as follows. By [Lemma 1](#), we see that $Q = c^{-2}Q_c(c^{-1} \cdot)$ is a symmetric-decreasing ground state for $\mathcal{E}(\psi)$ subject to $\int |\psi|^2 = N = c^{-1}$; and moreover the Lagrange multiplier $\mu(Q)$ for Q is found to be

$$\mu(Q) = c^{-2}\mu(Q_c). \quad (6-2)$$

Next, we consider the ground state energy $E(N)$ given by (1-2) for $0 < N < c_0^{-1}$. From [[Lieb and Yau 1987](#); [Fröhlich et al. 2007b](#)] we know that $E(N)$ is strictly concave. Hence $E'(N)$ exists for all $N \in (0, c_0^{-1}) \setminus \Sigma$, where Σ is some countable set, and we readily find that

$$E'(N) = -\mu(Q), \quad \text{for } N \in (0, c_0^{-1}) \setminus \Sigma. \quad (6-3)$$

Therefore the left-hand side of (6-2) only depends on $N = c^{-1}$ except when $N \in \Sigma$, which proves (6-1) with the countable set $\Xi = \{c : c > c_0 \text{ and } c^{-1} \in \Sigma\}$.

Suppose $\{c_n\}_{n=1}^\infty$ is a sequence with such that $c_n \rightarrow \infty$ and values in $c_n \in (c_0, \infty) \setminus \Xi$. Correspondingly, let $\{Q_{c_n}\}_{n=1}^\infty$ be a sequence of symmetric-decreasing ground states for $\mathcal{E}_c(\psi)$ with $\int |Q_{c_n}|^2 = 1$ for all $n \geq 1$. By [Proposition 1](#), for any such sequence $\{Q_{c_n}\}$, we have that Q_{c_n} and its corresponding Lagrange multipliers μ_{c_n} satisfy the assumption of [Proposition 3](#), provided that $n \gg 1$. By the local uniqueness result stated in [Proposition 3](#) and the fact μ_{c_n} only depends on c_n , we conclude that the symmetric-decreasing ground state Q_c for $\mathcal{E}_c(\psi)$ subject to $\int |\psi|^2 = 1$ is unique, provided that $c \in (c_0, \infty) \setminus \Xi$ holds, where $c_0 \gg 1$ is sufficiently large and Ξ is some countable set.

Finally, by [Lemma 1](#), we deduce uniqueness of symmetric-decreasing ground states Q for $\mathcal{E}(\psi)$ subject to $\int |\psi|^2 = N$, provided that $N \in (0, N_0) \setminus \Sigma$ holds, where $N_0 = c_0^{-1} \ll 1$ is sufficiently small and Σ denotes some countable set. \square

7. Proof of Theorems 3 and 4

We first prove [Theorem 4](#). By rescaling $Q_\infty(r) \mapsto aQ_\infty(br)$ with suitable $a > 0$ and $b > 0$, we can assume without loss of generality that $Q_\infty \in H_r^1(\mathbb{R}^3)$ satisfies the normalized equation

$$-\Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2) Q_\infty = -Q_\infty. \quad (7-1)$$

To complete the proof of [Theorem 4](#), it suffices to prove the following result.

Proposition 4. *Let $Q_\infty \in H_r^1(\mathbb{R}^3)$ be the unique radial and positive solution to [Equation \(7-1\)](#). Then the linearized operator L_+ given by*

$$L_+\zeta = -\Delta\zeta + \zeta - (|x|^{-1} * |Q_\infty|^2)\zeta - 2Q_\infty(|x|^{-1} * (Q_\infty\zeta)),$$

acting on $L^2(\mathbb{R}^3)$ with domain $H^1(\mathbb{R}^3)$, has the kernel

$$\ker L_+ = \text{span} \{ \partial_{x_1} Q_\infty, \partial_{x_2} Q_\infty, \partial_{x_3} Q_\infty \}.$$

Remark. For linearized operators L_+ arising from ground states Q for NLS with local nonlinearities, it is a well-known fact that $\ker L_+ = \{0\}$ when L_+ is restricted to radial functions implies that $\ker L_+$ is spanned by $\{\partial_{x_i} Q\}_{i=1}^3$.

The proof, however, involves some Sturm–Liouville theory which is not applicable to L_+ given above, due to the presence of the nonlocal term. (Also, recall that Newton’s theorem is not at our disposal, since we do not restrict ourselves to radial functions anymore.) To overcome this difficulty, we have to develop Perron–Frobenius-type arguments for the action of L_+ with respect to decomposition into spherical harmonics.

7.1. Proof of [Proposition 4](#). Since $Q_\infty(r)$ and $|x|^{-1}$ are radial functions, the operator L_+ commutes with rotations in \mathbb{R}^3 ; that is, we have that $(L_+\zeta(R\cdot))(x) = (L_+\zeta)(Rx)$ for all $R \in O(3)$. Therefore, we decompose any $\zeta \in L^2(\mathbb{R}^3)$ using spherical harmonics according to

$$\zeta(x) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} f_{\ell m}(r) Y_{\ell m}(\Omega), \quad (7-2)$$

where $x = r\Omega$ with $r = |x|$ and $\Omega \in \mathbb{S}^2$. This gives us the direct decomposition

$$L^2(\mathbb{R}^3) = \bigoplus_{\ell=0}^{\infty} \mathcal{H}_{(\ell)}, \quad (7-3)$$

so that L_+ acts invariantly on each

$$\mathcal{H}_{(\ell)} = L^2(\mathbb{R}_+, r^2 dr) \otimes \mathcal{Y}_{(\ell)}. \quad (7-4)$$

Here $\mathcal{Y}_{(\ell)} = \text{span} \{ Y_{\ell m} \}_{m=-\ell}^{+\ell}$ denotes the $(2\ell+1)$ -dimensional eigenspace corresponding to the eigenvalue $\kappa_\ell = -\ell(\ell+1)$ of the spherical Laplacian $\Delta_{\mathbb{S}^2}$ acting on $L^2(\mathbb{S}^2)$.

Let us now find an explicit formula for the action of L_+ on each $\mathcal{H}_{(\ell)}$. To this end, we recall the well-known the fact that

$$-\Delta = -\partial_r^2 - \frac{2}{r}\partial_r + \frac{\ell(\ell+1)}{r^2} \quad \text{on } \mathcal{H}_{(\ell)}, \quad (7-5)$$

as well as the multipole expansion

$$\frac{1}{|x-x'|} = 4\pi \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell} \frac{1}{2\ell+1} \frac{r_{<}^{\ell}}{r_{>}^{\ell+1}} Y_{\ell m}(\Omega) Y_{\ell m}^*(\Omega'), \quad (7-6)$$

where $r_{<} = \min(|x|, |x'|)$ and $r_{>} = \max(|x|, |x'|)$. An elementary calculation leads to the following equivalence: We have that $L_+\zeta = 0$ if and only if

$$L_{+,\ell} f_{\ell m} = 0, \quad \text{for } \ell = 0, 1, 2, \dots \text{ and } m = -\ell, \dots, +\ell, \quad (7-7)$$

with ζ given by (7-2). Here the operator $L_{+,\ell}$ acting on $L^2(\mathbb{R}_+, r^2 dr)$ is (formally) given by

$$(L_{+,\ell} f)(r) = -f''(r) - \frac{2}{r} f'(r) + \frac{\ell(\ell+1)}{r^2} f(r) + V(r)f(r) + (W_{(\ell)} f)(r), \quad (7-8)$$

with the local potential

$$V(r) = -(|x|^{-1} * |Q_{\infty}|^2)(r), \quad (7-9)$$

and the nonlocal linear operator

$$(W_{(\ell)} f)(r) = -\frac{8\pi}{2\ell+1} Q_{\infty}(r) \int_0^{\infty} \frac{r_{<}^{\ell}}{r_{>}^{\ell+1}} Q_{\infty}(s) f(s) s^2 ds, \quad (7-10)$$

where $r_{<} = \min(r, s)$ and $r_{>} = \max(r, s)$.

To prove Proposition 4, it suffices to assume henceforth that $\ell \geq 1$ holds, since $L_{+,0} f = 0$ implies that $f \equiv 0$ holds, by Proposition 2 above. Hence any nontrivial elements in the kernel of L_+ can only belong to $\mathcal{H}_{(\ell)}$ with $\ell \geq 1$. Before we proceed, we show that each $L_{+,\ell}$ enjoys a Perron–Frobenius property as follows.

Lemma 7. *For each $\ell \geq 1$, the operator $L_{+,\ell}$ is essentially self-adjoint on $C_0^{\infty}(\mathbb{R}_+) \subset L^2(\mathbb{R}_+, r^2 dr)$ and bounded below. Moreover, each $L_{+,\ell}$ has the Perron–Frobenius property. That is, if $e_{0,(\ell)}$ denotes the lowest eigenvalue of $L_{+,\ell}$, then $e_{0,(\ell)}$ is simple and the corresponding eigenfunction $\phi_{0,(\ell)}(r) > 0$ is strictly positive.*

Remarks. (1) We have indeed the lower bound $L_{+,\ell} \geq 0$ for all $\ell \geq 1$. This follows from $\mathcal{H}_{(\ell)} \perp Q_{\infty}$ for $\ell \geq 1$ and the fact that $L_+|_{Q_{\infty}^{\perp}} \geq 0$, which can be proven in the same way as for ground states for local NLS; see, for example, [Chang et al. 2007; Weinstein 1985].

(2) It is easy to see that $L_{+,\ell}$ has in fact infinitely many eigenvalues between 0 and 1. Indeed, the lower bound $Q_{\infty}(r) \geq B_{\varepsilon} e^{-(1+\varepsilon)r}$ (see the proof of Lemma 5) leads, by using Newton’s theorem, to the upper bound $V(r) \leq -\alpha r^{-1}$ with some $\alpha > 0$. Furthermore, one finds that $\langle f, W^{(\ell)} f \rangle < 0$ for $f \neq 0$. Hence, we conclude

$$L_{+,\ell} \leq -\partial_r^2 - \frac{2}{r} \partial_r + 1 + \frac{\ell(\ell+1)}{r^2} - \frac{\alpha}{r}$$

on $L^2(\mathbb{R}_+, r^2 dr)$. From the well-known spectral properties of the hydrogen atom Hamiltonian, we infer that the operator on the right has infinitely many eigenvalues below 1, and so does $L_{+,\ell}$ by the min-max principle.

Proof of Lemma 7. Since $Q_\infty(r)$ is exponentially decaying, it is straightforward to verify that $W_{(\ell)}$ is a bounded operator. Also, we have that $V \in L^\infty$ holds. Thus $L_{+,(\ell)}$ is bounded below (see also the remark following Lemma 7). Furthermore, it is well-known that

$$-\Delta_{(\ell)} = -\partial_r^2 - \frac{2}{r}\partial_r + \frac{\ell(\ell+1)}{r^2} \quad (7-11)$$

is essentially self-adjoint on $C_0^\infty(\mathbb{R}_+)$ provided that $\ell \geq 1$. In fact, this follows from [Reed and Simon 1980, Theorem X.10 and Example 4] which shows that $-\partial_r^2 - (2/r)\partial_r + \ell(\ell+1)/r^2$ is essentially self-adjoint on $C_0^\infty(\mathbb{R}_+)$ if $\ell(\ell+1)/r^2 \geq 3/4r^2$. Furthermore, by the Kato–Rellich theorem and the fact that V and $W_{(\ell)}$ are bounded and self-adjoint, we deduce that $L_{+,(\ell)} = -\Delta_{(\ell)} + V + W_{(\ell)}$ is essentially self-adjoint on $C_0^\infty(\mathbb{R}_+)$ as well.

The Perron–Frobenius property of $L_{+,(\ell)}$ can be shown as follows. First, we consider the kinetic energy part in $L_{+,(\ell)}$, where we find that

$$e^{t\Delta_{(\ell)}} \text{ is positivity improving on } L^2(\mathbb{R}_+, r^2 dr) \text{ for all } t > 0. \quad (7-12)$$

(Recall that, by definition, this means that $e^{t\Delta_{(\ell)}} f > 0$ when $f \geq 0$ with $f \not\equiv 0$.) Indeed, an argument given in Appendix B shows that the integral kernel of $e^{t\Delta_{(\ell)}}$ is strictly positive:

$$e^{t\Delta_{(\ell)}}(r, s) = \frac{1}{2t} \sqrt{\frac{1}{rs}} e^{-\frac{r^2+s^2}{4t}} I_{\ell+1/2}\left(\frac{rs}{2t}\right) > 0, \quad \text{for } r, s > 0. \quad (7-13)$$

Here $I_k(z)$ denotes the modified Bessel function of the first kind of order k . For later use, we record that (7-12) and the formula (by functional calculus)

$$(-\Delta_{(\ell)} + \mu)^{-1} = \int_0^\infty e^{-t\mu} e^{t\Delta_{(\ell)}} dt, \quad \text{for } \mu > 0, \quad (7-14)$$

immediately show that

$$(-\Delta_{(\ell)} + \mu)^{-1} \text{ is positivity improving on } L^2(\mathbb{R}_+, r^2 dr) \text{ for all } \mu > 0. \quad (7-15)$$

Next, let $A_{(\ell)}$ denote the bounded self-adjoint operator

$$A_{(\ell)} = V + W_{(\ell)}, \quad (7-16)$$

where V and $W_{(\ell)}$ are defined in (7-9) and (7-10), respectively. Note that $A_{(\ell)}$ is nonlocal. Using that $Q_\infty(r)$ is strictly positive, we readily find that

$$-A_{(\ell)} \text{ is positivity improving on } L^2(\mathbb{R}_+, r^2 dr). \quad (7-17)$$

This leads to the following auxiliary result.

Lemma 8. *For $\mu \gg 1$, the resolvent*

$$(L_{+,(\ell)} + \mu)^{-1} = (-\Delta_{(\ell)} + A_{(\ell)} + \mu)^{-1}$$

is positivity improving on $L^2(\mathbb{R}_+, r^2 dr)$.

Proof. For $\mu \gg 1$, we have

$$\frac{1}{L_{+, (\ell)} + \mu} = \frac{1}{-\Delta_{(\ell)} + \mu} \frac{1}{1 + A_{(\ell)}(-\Delta_{(\ell)} + \mu)^{-1}}.$$

Since $A_{(\ell)}$ is bounded, we conclude that $\|A_{(\ell)}(-\Delta_{(\ell)} + \mu)^{-1}\|_{L^2 \rightarrow L^2} < 1$ for $\mu \gg 1$. Thus a Neumann expansion yields

$$\frac{1}{L_{+, (\ell)} + \mu} = \frac{1}{-\Delta_{(\ell)} + \mu} \sum_{\nu=0}^{\infty} (-A_{(\ell)}(-\Delta_{(\ell)} + \mu)^{-1})^{\nu}, \quad (7-18)$$

provided that $\mu \gg 1$. Next, we recall from (7-15) that $(-\Delta_{(\ell)} + \mu)^{-1}$ is positivity improving. By this fact and (7-17), we deduce from (7-18) that $(L_{+, (\ell)} + \mu)^{-1}$ must be positivity improving for $\mu \gg 1$. This completes the proof of Lemma 8. \square

We now return to the proof of Lemma 7, which we complete as follows. Let $\ell \geq 1$ be fixed and suppose $e_{0, (\ell)} = \inf \sigma(L_{+, (\ell)})$ is the lowest eigenvalue. Furthermore, we choose $\mu \gg 1$ such that, by Lemma 8,

$$B = (L_{+, (\ell)} + \mu)^{-1} \quad (7-19)$$

is positivity improving on $L^2(\mathbb{R}_+, r^2 dr)$. Clearly, the operator B is bounded and self-adjoint, and its largest eigenvalue $\lambda_0 = \sup \sigma(B)$ is given by $\lambda_0 = (e_{(\ell), 0} + \mu)^{-1}$. Also, the corresponding eigenspaces of $L_{+, (\ell)}$ and B coincide. Since B is positivity improving (and hence ergodic), we can invoke [Reed and Simon 1978, Theorem XIII.43] to conclude that λ_0 is simple and that the corresponding eigenfunction $\phi_{(\ell), 0}(r)$ is strictly positive on \mathbb{R}_+ . This proof of Lemma 7 is therefore complete. \square

Let us now come back to the proof of Proposition 4, stating that $\ker L_+$ is spanned by $\{\partial_{x_i} Q_{\infty}\}_{i=1}^3$. By differentiating the nonlinear equation satisfied by Q_{∞} , we readily obtain that $L_+ \partial_{x_i} Q_{\infty} = 0$ for $i = 1, 2, 3$. Since $\partial_{x_i} Q_{\infty}(r) = Q'_{\infty}(r)(x_i/r) \in \mathcal{H}_{(1)}$, this show that

$$L_{+, (1)} Q'_{\infty} = 0. \quad (7-20)$$

Furthermore, by monotonicity of $Q_{\infty}(r)$, we have that $Q'_{\infty}(r) \leq 0$. Since $L_{+, (1)}$ is self-adjoint and Q'_{∞} is an eigenfunction that does not change its sign, Lemma 7 shows that in fact $Q'_{\infty}(r) = -\phi_{0, (1)}(r)$ holds, where $\phi_{0, (1)} > 0$ is the strictly positive ground state of $L_{+, (1)}$, with $e_{0, (1)} = 0$ being its corresponding eigenvalue. Therefore any $\zeta \in \mathcal{H}_{(1)}$ such that $L_+ \zeta = 0$ must be some linear combination of $\{\partial_{x_i} Q_{\infty}\}_{i=1}^3$.

To complete the proof of Proposition 4, we now claim that

$$L_{+, (\ell)} > 0, \quad \text{for } \ell \geq 2, \quad (7-21)$$

which in particular shows that $L_+ \zeta = 0$ with $\zeta \in \mathcal{H}_{(\ell)}$ for some $\ell \geq 2$ implies that $\zeta \equiv 0$. To prove (7-21), let $\ell \geq 2$ be fixed and set

$$e_{0, (\ell)} = \inf \sigma(L_{+, (\ell)}). \quad (7-22)$$

Indeed, by the remark following Lemma 7, we know that $e_{0, (\ell)} < 1$ is attained. (If $e_{0, (\ell)}$ was not attained, then $e_{0, (\ell)} = \inf \sigma_{\text{ess}}(L_{+, (\ell)}) = 1$ and (7-21) follows immediately.) By Lemma 7, the eigenvalue $e_{0, (\ell)}$ is simple and its corresponding eigenfunction $\phi_{0, (\ell)}(r) > 0$ is strictly positive. Next, we notice that

$$e_0 = \langle \phi_{0, (\ell)}, L_{+, (\ell)} \phi_{0, (\ell)} \rangle = \langle \phi_{0, (\ell)}, L_{+, (1)} \phi_{0, (\ell)} \rangle + K_{(\ell)}, \quad (7-23)$$

where

$$K_{(\ell)} = \int_0^\infty \frac{(\ell(\ell+1)-2)}{r^2} \phi_{0,(\ell)}(r)^2 r^2 dr + 8\pi \int_0^\infty \int_0^\infty Q_\infty(r) \phi_{0,(\ell)}(r) \left(\frac{1}{3} \frac{r_{<}}{r_{>}^2} - \frac{1}{2\ell+1} \frac{r_{<}^\ell}{r_{>}^{\ell+1}} \right) Q_\infty(s) \phi_{0,(\ell)}(s) r^2 s^2 dr ds,$$

with $r_{<} = \min(r, s)$ and $r_{>} = \max(r, s)$. Using the strict positivity of $Q_\infty(r)$ and $\phi_{0,(\ell)}(r)$, we see that $K_{(\ell)} > 0$ holds because of $\ell \geq 2$ and $(r_{<}/r_{>}) \leq 1$. Moreover, we recall from the preceding discussion that $L_{+, (1)} \geq e_{0, (1)} = 0$. Therefore, by (7-23),

$$e_{0, (\ell)} \geq K_{(\ell)} > 0, \quad \text{for all } \ell \geq 2, \quad (7-24)$$

which proves (7-21), completing the proof of Proposition 4, whence the proof of Theorem 4 follows. \square

7.2. Proof of Theorem 3. As in the proof of Theorem 2 above, it is convenient to fix $N > 0$ and to consider symmetric-decreasing ground state $Q_c \in H_r^1(\mathbb{R}^3)$ minimizing $\mathcal{E}_c(\psi)$ with $\int |Q_c|^2 = N$, where we take $c > 0$ sufficiently large. In what follows, let μ_c denote the Lagrange multiplier associated to Q_c . (It is possible that μ_c depends on Q_c and not just on c .)

Recall from Proposition 1 that

$$\|Q_c - Q_\infty\|_{H^1} \leq \delta_1 \quad \text{and} \quad |-\mu_c - mc^2 + \lambda| \leq \delta_2, \quad (7-25)$$

where $\delta_1 \rightarrow 0$ and $\delta_2 \rightarrow 0$ as $c \rightarrow \infty$. Here $Q_\infty \in H_r^1(\mathbb{R}^3)$ is the unique radial positive solution to (3-3) with $\int |Q_\infty|^2 = N$, where $\lambda > 0$ is determined through Q_∞ . By Theorem 4, the linear operator L_+ given by

$$L_+ \xi = -\frac{1}{2m} \Delta + \lambda - (|x|^{-1} * |Q_\infty|^2) \xi - 2Q_\infty (|x|^{-1} * (Q_\infty \xi)) \quad (7-26)$$

has the kernel

$$\ker L_+ = \text{span} \{ \partial_{x_1} Q_\infty, \partial_{x_2} Q_\infty, \partial_{x_3} Q_\infty \}. \quad (7-27)$$

Next, let $L_{+,c}$ denote the linear operators defined as

$$L_{+,c} \xi = \sqrt{-c^2 \Delta + m^2 c^4} \xi + \mu_c \xi - (|x|^{-1} * |Q_c|^2) \xi - 2Q_c (|x|^{-1} * (Q_c \xi)). \quad (7-28)$$

Again, upon differentiating the Euler–Lagrange equation satisfied by Q_c , we see that $L_{+,c} \partial_{x_i} Q_c = 0$ for $i = 1, 2, 3$. Hence

$$\text{span} \{ \partial_{x_1} Q_c, \partial_{x_2} Q_c, \partial_{x_3} Q_c \} \subseteq \ker L_{+,c}. \quad (7-29)$$

By the following perturbation argument, we show that in fact equality holds for $c \gg 1$. By standard arguments, we see that $0 \in \sigma(L_+)$ is an isolated eigenvalue. Thus we can construct the Riesz projection P_0 onto $\ker L_+$ by

$$P_0 = \frac{1}{2\pi i} \oint_{\Gamma_r} (L_+ - z)^{-1} dz, \quad (7-30)$$

where the curve Γ_r parametrizes the circle $\{z \in \mathbb{C} : |z| = r\}$. Here $r > 0$ is chosen sufficiently small such that 0 is the only eigenvalue of L_+ inside $|z| \leq r$. Next, we claim that the projection

$$P_{0,c} = \frac{1}{2\pi i} \oint_{\Gamma_r} (L_{+,c} - z)^{-1} dz \quad (7-31)$$

exists for $c \gg 1$ and satisfies

$$\|P_{0,c} - P_0\|_{L^2 \rightarrow L^2} \rightarrow 0 \quad \text{as } c \rightarrow \infty. \quad (7-32)$$

Indeed, by using (7-25) and similar arguments as in the proof of Proposition 3 (see, for example, the resolvent estimate (5-13)), we conclude that

$$\|(L_{+,c} - z)^{-1}\|_{L^2 \rightarrow L^2} \leq C \|(L_+ - z)^{-1}\|_{L^2 \rightarrow L^2}, \quad (7-33)$$

for all $c \gg 1$ and $z \in \Gamma_r$, where $C > 0$ is some constant. Furthermore, we have

$$\|(L_{+,c} - z)^{-1} - (L_+ - z)^{-1}\|_{L^2 \rightarrow L^2} \rightarrow 0 \quad \text{as } c \rightarrow \infty, \quad (7-34)$$

for all $z \in \Gamma_r$. This shows that $P_{0,c}$ exists for $c \gg 1$ and that (7-32) holds. Since $\text{rank } P_0 = 3$ and the rank of $P_{0,c}$ remains constant for $c \gg 1$, by (7-32), we infer that $P_{0,c}$ has at most 3 eigenvalues (counted with their multiplicity) inside $|z| \leq r$, provided that $c \gg 1$. In particular, we conclude that $\dim \ker L_{+,c} \leq 3$ for $c \gg 1$. Therefore equality must hold in (7-29) whenever $c \gg 1$.

Thus we have found that $L_{+,c}$ has the desired kernel property if $c \gg 1$. By a rescaling argument formulated in Lemma 1, we conclude the analogous statement for the linear operator L_+ arising from the unique symmetric-decreasing ground state Q minimizing $\mathcal{E}(\psi)$ subject to $\int |\psi|^2 = N$ with $N \ll 1$. The proof of Theorem 3 is now complete. \square

Appendix A. Uniqueness of Q_∞

Suppose that $Q_\infty \in H^1(\mathbb{R}^3)$ solves

$$-\frac{1}{2m} \Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2) Q_\infty = -\lambda Q_\infty, \quad (A-1)$$

with $m > 0$ and $\lambda > 0$ given. By rescaling $Q_\infty(r) \mapsto a Q_\infty(br)$ with suitable $a > 0$ and $b > 0$, we can consider without loss of generality solutions $Q_\infty \in H^1(\mathbb{R}^3)$ to the “normalized” equation

$$-\Delta Q_\infty - (|x|^{-1} * |Q_\infty|^2) Q_\infty = -Q_\infty. \quad (A-2)$$

The following result is due to [Lieb 1977]; see also [Tod and Moroz 1999]. Here we provide a partly different proof, which is directly based on a comparison argument.

Lemma 9. *Equation (A-2) has a unique radial, nonnegative solution $Q \in H_r^1(\mathbb{R}^3)$ with $Q \not\equiv 0$. Moreover, we have that $Q(r)$ is in fact strictly positive.*

Proof. Existence of a nonnegative, nontrivial solution $Q_\infty \in H_r^1(\mathbb{R}^3)$ of (A-2) follows from variational arguments; see [Lieb 1977].

To prove that any nonnegative $Q \in H^1(\mathbb{R}^3)$, with $Q \not\equiv 0$, solving (A-2) is strictly positive, we can simply argue as follows. We rewrite (A-2) as

$$Q(x) = ((-\Delta + 1)^{-1}(VQ))(x) = \frac{1}{4\pi} \int_{\mathbb{R}^3} \frac{e^{-|x-y|}}{|x-y|} V(y) Q(y) dy \quad (A-3)$$

with $V = |x|^{-1} * |Q|^2$. Since $V \geq 0$ and $Q \geq 0$ (with $V \not\equiv 0$ and $Q \not\equiv 0$), Equation (A-3) shows that Q is strictly positive.

Let us now prove the claimed uniqueness. Suppose $Q \in H_r^1(\mathbb{R}^3)$, with $Q \not\equiv 0$, is a solution to (A-2). Using Newton's theorem, we find that $Q(r)$ solves (after a suitable rescaling $Q(r) \mapsto a^2 Q(ar)$ for some $a > 0$; see [Lieb 1977]) the initial-value problem

$$-v''(r) - \frac{2}{r}v'(r) - v(r) + \left(\int_0^r K(r,s)v(s)^2 ds \right)v(r) = 0, \quad v(0) = v_0, \quad v'(0) = 0, \quad (\text{A-4})$$

with $v_0 = Q(0) \in \mathbb{R}$. (Recall that $K(r,s) \geq 0$ is given by (4-6) above.) By standard fixed point arguments, we deduce that (A-4) has a unique local C^2 -solution for given initial data $v(0) \in \mathbb{R}$ and $v'(0) = 0$, and $v(r)$ exists up to some maximal radius $R \in (0, \infty]$.

Suppose now that $Q \in H_r^1(\mathbb{R}^3)$ and $\tilde{Q} \in H_r^1(\mathbb{R}^3)$ are two radial, nonnegative (and nontrivial) solutions to (A-2) with $Q \not\equiv \tilde{Q}$. From the preceding discussion we know that Q and \tilde{Q} are in fact strictly positive, and (after appropriate rescaling) both satisfy (A-4) with $v_0 = Q(0) > 0$ and $v_0 = \tilde{Q}(0) > 0$, respectively. By uniqueness for (A-4), we conclude that $Q(0) \neq \tilde{Q}(0)$ holds, since otherwise $Q \equiv \tilde{Q}$. Therefore, we can henceforth assume that

$$\tilde{Q}(0) > Q(0). \quad (\text{A-5})$$

Next, we notice that a calculation (similar to the one in the proof of Lemma 5) yields the integrated ‘‘Wronskian-type’’ identity

$$r^2(Q(r)\tilde{Q}'(r) - Q'(r)\tilde{Q}(r)) = \int_0^r s^2 Q(s)\tilde{Q}(s)(\tilde{V}(s) - V(s)) ds. \quad (\text{A-6})$$

Here,

$$V(r) = \int_0^r K(r,s)Q(s)^2 ds \quad \text{and} \quad \tilde{V}(r) = \int_0^r K(r,s)\tilde{Q}(s)^2 ds. \quad (\text{A-7})$$

By continuity and (A-5), we have $\tilde{Q}(r) > Q(r)$ at least initially for $r \geq 0$. Next, we conclude, by (A-6), that in fact

$$\tilde{Q}(r) > Q(r), \quad \text{for all } r \geq 0. \quad (\text{A-8})$$

To see this, suppose on the contrary that $\tilde{Q}(r) > 0$ intersects $Q(r) > 0$ for the first time at $r = r_* > 0$, say. Then the left-hand side of (A-6) is found to be nonnegative at $r = r_*$, whereas the right-hand side must be strictly positive at $r = r_*$ since $\tilde{V}(r) > V(r)$ on $(0, r_*)$. This contradiction shows that (A-8) holds.

Finally, we show that (A-8) leads to a contradiction (along the lines of [Lieb 1977]) as follows. To this end, we consider the Schrödinger operators

$$H = -\Delta + V \quad \text{and} \quad \tilde{H} = -\Delta + \tilde{V}, \quad (\text{A-9})$$

so that $HQ = Q$ and $\tilde{H}\tilde{Q} = \tilde{Q}$. By standard theory of Schrödinger operators, we conclude that Q and \tilde{Q} are (up to a normalization factor) the unique positive ground states (with eigenvalue $e = 1$) for H and \tilde{H} , respectively. Therefore,

$$\langle \phi, H\phi \rangle \geq \|\phi\|_{L^2}^2 \quad \text{and} \quad \langle \phi, \tilde{H}\phi \rangle \geq \|\phi\|_{L^2}^2, \quad \text{for } \phi \in H^1(\mathbb{R}^3), \quad (\text{A-10})$$

where equality holds if and only if $\phi = \lambda Q$ or $\phi = \lambda \tilde{Q}$ for some $\lambda \in \mathbb{C}$, respectively.

Going back to (A-8), we find that $\tilde{V}(r) > V(r)$ for all $r > 0$, which leads to

$$\|\tilde{Q}\|_{L^2}^2 \leq \langle \tilde{Q}, H\tilde{Q} \rangle = \langle \tilde{Q}, \tilde{H}\tilde{Q} \rangle - \langle \tilde{Q}, (\tilde{V} - V)\tilde{Q} \rangle = \|\tilde{Q}\|_{L^2}^2 - \delta,$$

for some $\delta > 0$, which is a contradiction.

Hence (A-2) does not admit two different radial and nonnegative (and nontrivial) solutions $Q \in H_r^1(\mathbb{R}^3)$ and $\tilde{Q} \in H_r^1(\mathbb{R}^3)$. \square

Appendix B. Decomposition of $e^{t\Delta}$ using spherical harmonics

Recall the explicit formula for the heat kernel of the Laplacian Δ on \mathbb{R}^3 :

$$e^{t\Delta}(x, y) = \frac{1}{(4\pi t)^{3/2}} e^{-|x-y|^2/(4t)} = \frac{1}{(4\pi t)^{3/2}} e^{-(x^2+y^2)/(4t)} e^{(x \cdot y)/(2t)}. \quad (\text{B-1})$$

Moreover, we have the well-known identity

$$e^{ax \cdot y} = 4\pi \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell} i_{\ell}(a|x||y|) Y_{\ell m}(\Omega) Y_{\ell m}^*(\Omega') \quad (\text{B-2})$$

for $a > 0$, $x = |x|\Omega$ and $y = |y|\Omega'$ where $\Omega, \Omega' \in \mathbb{S}^2$. Here

$$i_{\ell}(z) = \sqrt{\frac{\pi}{2z}} I_{\ell+1/2}(z) \quad (\text{B-3})$$

is the modified spherical Bessel function of the first kind of order ℓ ; whereas $I_k(z)$ denotes the modified Bessel function of the first kind of order k .

Let $\Delta_{(\ell)}$ denote the restriction of Δ acting on $\mathcal{H}_{(\ell)}$ (that is, the space of $L^2(\mathbb{R}^3)$ functions whose “angular momentum” is $\ell \geq 0$). From (B-1) and (B-2) we deduce that the integral kernel of $e^{t\Delta_{(\ell)}}$ acting on $L^2(\mathbb{R}_+, r^2 dr)$ is given by

$$e^{t\Delta_{(\ell)}}(r, s) = \frac{1}{2t} \sqrt{\frac{1}{rs}} e^{-(r^2+s^2)/(4t)} I_{\ell+1/2}\left(\frac{rs}{2t}\right). \quad (\text{B-4})$$

An explicit integral representation for $I_k(z)$ shows that $I_{\ell+1/2}(z) > 0$ for all $z > 0$ and $\ell \geq 0$.

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