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SLOSHING FREQUENCIES IN A HALF-SPACE BY KELVIN INVERSION

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The frequencies of the free oscillations of an incompressible and inviscid fluid in a half-space with circular and strip-like aperture are determined by first mapping the half-space into a bounded domain by the Kelvin inversion. Upper bounds for the eigenfrequencies are then obtained by a Rayleigh-Ritz procedure. The results compare quite well with previous results, which were obtained by less elementary tools.

1. Introduction. The free oscillations of an incompressible, inviscid liquid in a container have been investigated extensively (see [1] and [6], and the references given there). In this paper we determine approximations to the sloshing frequencies for fluid in a covered half-space, where the free surface is a circular aperture. The problem represents the limiting case of free sloshing motions in arbitrary axially symmetric containers with the same free surface, and its solution furnishes universal upper bounds for the sloshing frequencies in these containers [11]. Recently, quite accurate results have been obtained in [4], including rigorous error bounds for the frequencies. Although the underlying geometry and the problem statement are quite simple, the solution with the error bounds turned out to be rather intricate.

We will present here an alternate method and show that quite acceptable upper bounds may be computed with considerably more modest tools. There are two reasons that account for the simplifications. First of all we will choose coordinate functions in the Rayleigh-Ritz procedure, which do not satisfy the natural boundary conditions of the variational formulation of the problem. This leads to considerable simplifications, but it precludes, on the other hand, the ready computation of lower bounds. Secondly, the complications arising from the fact that the domain is unbounded will be avoided by using a Kelvin inversion. The problem is solved for the image domain of a large spherical bowl, and only in the final result is the bowl radius assumed to tend to infinity.

The corresponding variational formulation can still be established without difficulty, although the problem is slightly more complicated after the Kelvin transformation (cf. eqs. (2.5) to (2.8) below). In this setting, the problem is solved by a Rayleigh-Ritz method. The matrix elements turn out to be rational numbers (with the

exception of a few logarithmic terms) and can therefore be computed easily and accurately.

The mathematical problem, both in its original form and after the Kelvin transformation, is stated in § 2. The next section describes in detail how the original differential equation and the boundary conditions transform under the Kelvin inversion. The § 4 establishes the variational principle, and § 5 deals with the choice of the coordinate functions in the Rayleigh-Ritz procedure. In § 6 it is indicated how the integrals, which represent the elements in the matrix eigenvalue problem, are evaluated.

The discussion of the limiting process from the large spherical bowl to a half-space is taken up in § 7. The next section comments on the two-dimensional sloshing problems in a half-space with a strip-like aperture, and the numerical results for both the axially symmetric and the plane case are given in the last section.

2. Problem statement. The sloshing motion of a fluid in a half-space with a circular free surface of radius a is described by the following eigenvalue problem for the velocity potential $f(x, y, z)$ [4], where the eigenvalue appears only in the boundary condition.

The Laplace equation is satisfied in the lower half-space $z < 0$

$$\nabla^2 f = 0 , \tag{2.1}$$

and the boundary conditions on the horizontal x-y-plane are

$$\frac{\partial f}{\partial z}(x, y, 0) = \lambda f \ (x, y, 0) \quad \text{for} \quad 0 \leqq x^2 + y^2 < a^2 \tag{2.2}$$

$$\frac{\partial f}{\partial z}(x, y, 0) = 0 \qquad \text{for} \quad a^2 < x^2 + y^2 < \infty . \tag{2.3}$$

The eigenvalues λ are related to the frequencies of oscillation ω and to the acceleration of gravity g by

$$\lambda = \omega^2/g .$$

In the approach taken in this paper, the eqs. (2.1) to (2.3) are satisfied inside a large spherical bowl and on the part of the $x-y$-plane which cuts the bowl, whereas on the spherical part of the bowl the usual boundary condition for a finite container is imposed (cf. Fig. 1), namely that the normal derivative of the velocity potential vanishes

$$\frac{\partial f}{\partial n} = 0 \qquad \text{on} \quad S_1 . \tag{2.4}$$

If the Kelvin inversion is applied, then the problem takes on the

following form, in which it is actually solved. Determine the set of eigenvalues λ for the eigenvalue problem for the function g:

$$\nabla^2 g = 0 \tag{2.5}$$

in the indented sphere C_3 of radius $a/2$ (see Fig. 1); the boundary condition

$$g + a\frac{\partial g}{\partial r} = 0 \tag{2.6}$$

is satisfied on the upper half-sphere, except on the indentation of radius ρ. On the indentation

$$\rho\frac{\partial g}{\partial n} = g \tag{2.7}$$

holds (n is the outer normal to the shaded domain). Furthermore,

$$(1 - \cos\theta)\left(g + a\frac{\partial g}{\partial r}\right) = 2\lambda a g \tag{2.8}$$

holds on the lower half of the sphere C_3.

The derivation of this form of the problem is carried out in the next section.

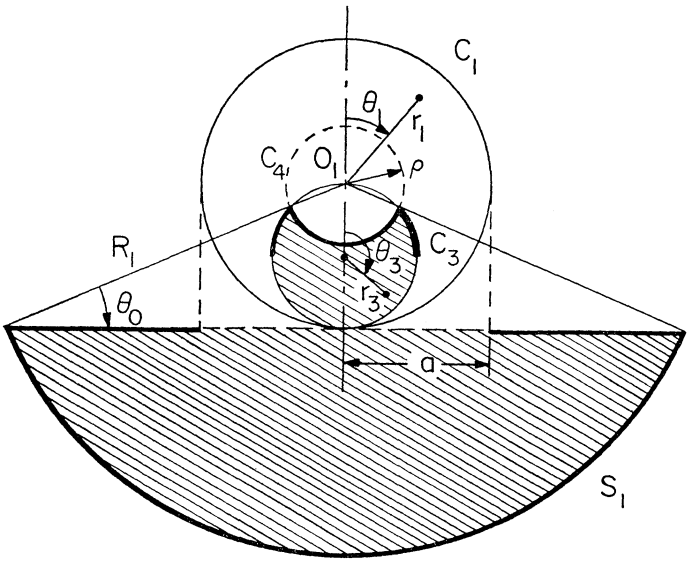

FIG. 1 The geometry of the Kelvin inversion.

A remark should be added about the singularity which occurs at the transition from the free surface to the covered part of the

spherical bowl, i.e., at $x^2 + y^2 = a^2$, $z = 0$. From the physical point of view we have to stipulate that the singularity be weaker than a source or a sink (cf. [10], p. 59 and p. 80). This is also the appropriate mathematical condition for the problem: as P. Henrici has shown in a striking example (cf. [4], p. 296), the spectrum of the problem changes its character drastically (there appear finite eigenvalues λ of infinite multiplicity), if the singularity is too strong. Since the singularity at the rim is a local phenomenon, the solution of the dock problem (waves in a half-space covered by a rigid half-plane) is applicable. This solution is given in [3] (and also used in [8], p. 323) in terms of the complex velocity potential: there is a logarithmic singularity at the rim. For the Rayleigh-Ritz procedure used below the exact nature of the sufficiently weak singularity need not be taken into account. The coordinate functions (see § 5) will probably represent the true eigenfunctions rather poorly near the rim; but it is known that the Rayleigh-Ritz method nevertheless furnishes good upper bounds for the eigenvalues. Under some circumstances it is advisable to isolate the singular behavior of a solution by including in the set of coordinate functions a function with the proper singularity. However, the computation of the matrix elements in § 6 would then become considerably more involved.

Incidentally, the solution of the dock problem ([3], [8]) has been used in [12] to obtain the asymptotic distribution of the eigenvalues for the problem considered here.

3. The Kelvin inversion on the sphere C_1. The original problem is first restated in spherical coordinates r_1, θ_1, φ_1 with the origin at the center O_1 of the sphere of inversion C_1. The x-y-plane is then described by

$$r_1 \cos\theta_1 + a = 0 \,, \tag{3.1}$$

and its normal unit vector in the z-direction by

$$\vec{n} = (\cos\theta_1, -\sin\theta_1, 0) \,.$$

Hence the boundary conditions, eqs. (2.2), (2.3), and (2.4), for the harmonic function

$$f_1(r_1, \theta_1, \varphi_1) = f(x, y, z)$$

become, on the plane (3.1),

$$\frac{\partial f_1}{\partial r_1}\cos\theta_1 - \frac{\sin\theta_1}{r_1}\frac{\partial f_1}{\partial \theta_1} = \lambda f_1 \qquad \text{for } \frac{3\pi}{4} < \theta_1 \leqq \pi$$

$$\frac{\partial f_1}{\partial r_1}\cos\theta_1 - \frac{\sin\theta_1}{r_1}\frac{\partial f_1}{\partial \theta_1} = 0 \qquad \text{for } \frac{\pi}{2} + \theta_0 \leqq \theta_1 < \frac{3\pi}{4},$$

and furthermore,

$$\frac{\partial f_1}{\partial r_1} = 0$$

on the large spherical bowl S_1 of radius R_1. For the meaning of the angle θ_0 see Fig. 1, where it appears greatly exaggerated, of course.

Next we map the lower half-space into the sphere C_3 by reciprocal radii, i.e., by the transformation

$$r_2 = \frac{a^2}{r_1}, \qquad \theta_2 = \theta_1, \qquad \varphi_2 = \varphi_1.$$

For the equation of the sphere C_3 we then obtain

$$a\cos\theta_2 + r_2 = 0 \qquad \text{for } \frac{\pi}{2} \leqq \theta_2 \leqq \pi$$

and for the spherical indentation C_4 with radius ρ

$$r_2 = \rho$$

where $\rho = a^2/R_1$.

If we transform f_1 in the usual manner by introducing ([9] p. 140)

$$f_2(r_2, \theta_2, \varphi_2) = \frac{a^2}{r_2} f_1\left(\frac{a^2}{r_2}, \theta_2, \varphi_2\right),$$

then f_2 is again a harmonic function

$$\nabla^2 f_2 = 0. \tag{3.2}$$

The boundary conditions on C_3 turn out to be

$$r_2\cos\theta_2 f_2 + r_2^2\cos\theta_2\frac{\partial f_2}{\partial r_2} + r_2\sin\theta_2\frac{\partial f_2}{\partial \theta_2} = -\lambda a^2 f_2 \tag{3.3}$$

$$\text{for } \frac{3\pi}{4} < \theta_2 \leqq \pi,$$

$$\cos\theta_2 f_2 + r_2\cos\theta_2\frac{\partial f_2}{\partial r_2} + \sin\theta_2\frac{\partial f_2}{\partial \theta_2} = 0 \tag{3.4}$$

$$\text{for } \frac{\pi}{2} + \theta_0 \leqq \theta_2 < \frac{3\pi}{4},$$

and finally on the indentation

$$r_2\frac{\partial f_2}{\partial r_2} + f_2 = 0 \qquad \text{for } r_2 = \rho. \tag{3.5}$$

Clearly, the proper coordinates to solve (3.2) to (3.4) are the spherical coordinates with the origin at the center of the sphere C_3 of radius $a/2$, i.e.,

$$r_3 \cos\theta_3 = r_2 \cos\theta_2 + \frac{a}{2}$$
$$r_3 \sin\theta_3 = r_2 \sin\theta_2$$
$$\varphi_3 = \varphi_2 .$$

On the surface C_3 the relations

$$\frac{\partial r_3}{\partial r_2} = \sin\left(\frac{\theta_3}{2}\right)$$
$$\frac{\partial r_3}{\partial \theta_2} = -a \sin\left(\frac{\theta_3}{2}\right) \cos\left(\frac{\theta_3}{2}\right)$$
$$\frac{\partial \theta_3}{\partial r_2} = \frac{2}{a} \cos\left(\frac{\theta_3}{2}\right)$$
$$\frac{\partial \theta_3}{\partial \theta_2} = 2 \sin^2\left(\frac{\theta_3}{2}\right)$$

hold. Hence, the boundary conditions (3.3), (3.4) transform, after some straightforward simplifications, to

$$(1 - \cos\theta_3)\left(f_3 + a\frac{\partial f_3}{\partial r_3}\right) = 2\lambda a f_3 \tag{3.6}$$

on the lower half-sphere, denoted by S_l:

$$r_3 = \frac{a}{2}, \quad \frac{\pi}{2} < \theta_3 \leqq \pi ,$$

and to

$$f_3 + a\frac{\partial f_3}{\partial r_3} = 0 \tag{3.7}$$

on the upper half-sphere:

$$r_3 = \frac{a}{2}, \quad \bar{\theta} \leqq \theta_3 < \frac{\pi}{2} ,$$

where $f_3(r_3, \theta_3, \varphi_3) = f_2(r_2, \theta_2, \varphi_2)$ satisfies

$$\nabla^2 f_3 = 0 \tag{3.8}$$

inside the indented sphere C_3. The angle $\bar{\theta} > 0$ excludes the indentation from C_3. The boundary condition (3.5) on the indentation is best left in the old coordinate system. Thus the problem statement

of § 2 is established, and the subscripts can be dropped for the spherical coordinates.

4. **The variational principle.** It is not difficult to find by inspection the variational principle which leads to the eigenvalue problem (2.5) to (2.8).

Let the functional $L[g]$ be defined as

$$L[g] = N[g]/D[g] \tag{4.1}$$

with

$$N[g] = \iiint_V |\nabla g|^2 \, dV + \frac{1}{a} \iint_S g^2 \, dS - \frac{1}{\rho} \iint_{C_4} g^2 \, dS \,, \tag{4.2}$$

$$D[g] = \iint_{S_1} \frac{2}{(1 - \cos\theta)} g^2 \, dS \,. \tag{4.3}$$

In $N[g]$ the volume integral is taken over the indented sphere, one surface integral is taken over the surface S of the sphere C_3, without the indentation, and the other surface integral over the indentation C_4. In $D[g]$ the surface integral extends over the lower half-sphere. We show now by the standard method of the calculus of variations that the stationary values of L are the eigenvalues λ of the problem (2.5) to (2.8):

$$\iiint_V \nabla g \cdot \nabla \, \delta g \, dV + \frac{1}{a} \iint_S g \, \delta g \, dS - \frac{1}{\rho} \iint_{C_4} g \, \delta g \, dS - \lambda \iint_{S_1} \frac{2}{(1 - \cos\theta)} g \, \delta g \, dS = 0 \,,$$

and using Green's theorem

$$-\iiint_V \delta g \, \nabla^2 g \, dV + \iint_S \delta g \frac{\partial g}{\partial n} \, dS + \iint_{C_4} \delta g \frac{\partial g}{\partial n} \, dS + \frac{1}{a} \iint_S \delta g \, g \, dS - \frac{1}{\rho} \iint_{C_4} \delta g \, g \, dS - \lambda \iint_{S_1} \frac{2}{(1 - \cos\theta)} \delta g \, g \, dS = 0 \,.$$

The differential equation and the boundary conditions (2.5) to (2.8) follow now at once. We notice, in particular, that the boundary conditions turn out to be natural boundary conditions. This enables us to choose as admissible function in the variational principle any continuous function with piecewise continuous first derivatives for which the integrals exist.

5. **The Rayleigh-Ritz method.** We now proceed to find the stationary values of (4.1) by the Rayleigh-Ritz method. Since the true

eigenfunctions of the problem are harmonic, it is desirable to choose as coordinate functions a complete system of harmonic functions (cf. [6] p. 240, [7] p. 96). An obvious choice for g is then

$$g = \sum_{n=m}^{n=N} \alpha_n f_n \cos m\varphi = \sum_{n=m}^{n=N} \alpha_n \left(\frac{2r}{a}\right)^n P_n^m(\cos\theta) \cos m\varphi \,, \tag{5.1}$$

$$m = 1, 2, \cdots .$$

The functions P_n^m are the spherical harmonics, and the α_n's are the free coefficients in the Rayleigh-Ritz method.

For the sloshing modes without radial nodal lines, the above expression for g is not the best possible choice, because for $m = 0$ the original problem possesses the eigenvalue $\lambda = 0$ with a constant eigenfunction. Under the inversion, this function does not remain constant and should be included in the set of coordinate functions. Although the Rayleigh-Ritz method would still furnish upper bounds for the eigenvalues, the results would be considerably less accurate without this function, especially for small N. For this additional function the index $n = -1$ is chosen, in order not to disturb the standard notation for the Legendre polynomials P_n. By the inversion on the sphere one obtains

$$f_{-1} = \frac{a}{r_2} = \left(\frac{r^2}{a^2} + \frac{1}{4} - \frac{r}{a}\cos\theta\right)^{-1/2} \tag{5.2}$$

If $N \to \infty$, this coordinate function would of course be redundant. For $m = 0$ we now choose

$$g = \alpha_{-1} f_{-1} + \sum_{n=0}^{n=N} \alpha_n \left(\frac{2r}{a}\right)^n P_n(\cos\theta) \,. \tag{5.3}$$

Since the trigonometric functions are orthogonal, the problem is decomposed into its symmetry classes $m = 0, 1, 2, \cdots$, and each symmetry class can be investigated separately. This fact has already been taken into account in the coordinate functions introduced above. The factor $\cos m\varphi$ could of course be replaced by $\sin m\varphi$ without altering the subsequent steps, and it then follows that all eigenvalues for $m \geqq 1$ are double eigenvalues.

We are now ready to find the upper bounds for the functional $L[g]$ in eq. (4.1). In order to take advantage of the fact that all the coordinate functions are harmonic, the eq. (4.2) is transformed by Green's theorem

$$N[g] = \iint_S g\left(\frac{\partial g}{\partial r} + \frac{g}{a}\right) dS + \iint_{C_4} g\left(\frac{\partial g}{\partial n} - \frac{g}{\rho}\right) dS \,.$$

For each symmetry class m, eq. (5.1) leads to

$$L_m[g] = N_m[g]/D_m[g]$$

$$N_m[g] = \sum_{j,k} \alpha_j \alpha_k \left\{ \iint_S f_j \left(\frac{\partial f_k}{\partial r} + \frac{f_k}{a} \right) \cos^2 m\varphi \, dS \right.$$
$$\left. + \iint_{C_4} f_j \left(\frac{\partial f_k}{\partial n} - \frac{f_k}{\rho} \right) \cos^2 m\varphi \, dS \right\}$$

$$D_m[g] = \sum_{j,k} \alpha_j \alpha_k \iint_{S_l} f_j f_k \frac{2}{(1 - \cos\theta)} \cos^2 m\varphi \, dS \, .$$

If we minimize L with respect to the coefficients α ([7] p. 96), we obtain the algebraic eigenvalue problem of the form

$$\det(A - \lambda B) = 0 \, .$$

6. The computation of the matrix elements. The matrix elements are, after suppressing the common factors from the integration over φ

$$a_{jk} = \int_{\bar{\theta}}^{\pi} f_j \left(\frac{\partial f_k}{\partial r} + \frac{f_k}{a} \right) \frac{a^2}{4} \sin\theta \, d\theta$$
$$+ \int_{\pi/2+\theta_0}^{\pi} f_j \left(\frac{\partial f_k}{\partial n} - \frac{f_k}{\rho} \right) \rho^2 \sin\theta_2 \, d\theta_2 \tag{6.1}$$

$$b_{jk} = \int_{\pi/2}^{\pi} f_j f_k \frac{2}{(1 - \cos\theta)} \frac{a^2}{4} \sin\theta \, d\theta \, . \tag{6.2}$$

It is convenient to leave the second integral of a_{jk} in the spherical coordinates centered at O_1. In spite of their appearance, both matrices A and B are symmetrical.

The matrix elements a_{jk} and b_{jk} are evaluated under the assumption that $\theta_0 = 0$, and hence $\bar{\theta}_0 = 0, \rho = 0$. These assumptions will be justified in the next section. For all m, and $j \geqq m$, $k \geqq m$

$$f_k = \left(\frac{2r}{a} \right)^k P_k^m(\cos\theta) \, ,$$

and setting $\cos\theta = x$, the eqs. (6.1) and (6.2) simply become

$$a_{jk} = \frac{a}{4} \int_{-1}^{1} (2k+1) P_j^m(x) P_k^m(x) \, dx = \frac{a}{2} \frac{(j+m)!}{(j-m)!} \delta_{jk} \tag{6.3}$$

with the Kronecker symbol δ_{jk}(see [5], p. 116), and

$$b_{jk} = \frac{a^2}{4} \int_{-1}^{0} P_j^m(x) P_k^m(x) \frac{2}{(1-x)} \, dx \, . \tag{6.4}$$

For $m = 0$, there are the additional elements $a_{-1,k}$, a_{-1-1}, $b_{-1,k}$, $b_{-1,-1}$. Using eq. (5.2) for $r = a/2$, we obtain

$$b_{-1,k} = \frac{a^2}{4}\int_{-1}^{0} \frac{2\sqrt{2}}{(1-x)^{3/2}} P_k(x)\,dx\,, \qquad k = 0, 1, \cdots \tag{6.5}$$

$$b_{-1,-1} = \frac{a^2}{4}\int_{-1}^{0} \frac{4}{(1-x)^2}\,dx = \frac{a^2}{2}\,. \tag{6.6}$$

Both integrals in the elements $a_{k,-1}$ and $a_{-1,-1}$ vanish, since we find from eq. (5.2) that on C_3

$$\frac{\partial f_{-1}}{\partial r} + \frac{f_{-1}}{a} = 0 \tag{6.7}$$

and on the indentation C_4

$$\frac{\partial f_{-1}}{\partial n} - \frac{f_{-1}}{\rho} = 0\,. \tag{6.8}$$

As the matrix A is symmetrical, the elements $a_{-1,k}$ also vanish; indeed, in this case, the two integrals are equal in magnitude and of opposite sign.

It remains to evaluate the integrals in eqs. (6.4) and (6.5). If the Legendre functions P are expressed in terms of $(1 - x)$, as given in [5] p. 111, and the division is carried out, then the finite series can be integrated termwise. The results look cumbersome, if they are spelled out, but they are easily implemented recursively in a computer program.

7. **The limit of an infinitely large spherical bowl.** The integrals in eq. (6.1) have been evaluated for $\theta_0 = 0$, and hence $\bar{\theta} = 0$, $\rho = 0$. This is permissible, because the contribution of the indentation in the sphere C_3 to the matrix elements can be made arbitrarily small. All the coordinate functions, with the exception of f_{-1}, are regular in the neighborhood of the point O_1. By shrinking the radius ρ of the indentation C_4 to zero, we can make the contribution of the indentation to the matrix elements a_{jk} as small as we wish. In the elements $a_{k,-1}$ and $a_{-1,-1}$ the integrand vanishes even for finite ρ according to (6.7) and (6.8). This result is of course expected, as $\lambda = 0$ should remain an eigenvalue for the problem after the Kelvin inversion.

The roots λ of the characteristic polynomial are all nonnegative and depend continuously on its coefficients, i.e., on the elements of the matrices A and B. Therefore, if the indentation is ignored entirely, the Rayleigh-Ritz eigenvalues differ arbitrarily little from those for a small indentation, or, in terms of the original problem, from the Rayleigh-Ritz eigenvalues of the lower half-space bounded by a sphere of large radius R_1.

8. **The planar case.** For the case where the sloshing motion takes place in two dimensions, the same method can be applied. The free surface is now a strip of width $2a$ perpendicular to the plane in which the velocity vectors lie. Fig. 1 still depicts the situation, except that the inversion takes place on a circle rather than on a sphere. The sloshing frequencies now represent universal upper bounds for sloshing in a canal where the fluid motion is perpendicular to the canal axis.

The mathematical problem which describes the motion is, in terms of the velocity potential f, as follows:

$$\nabla^2 f = 0 \qquad \text{for } z < 0 \tag{8.1}$$

$$\frac{\partial f}{\partial z}(x, 0) = \lambda f(x, 0) \qquad \text{for } 0 < |x| < a \tag{8.2}$$

$$\frac{\partial f}{\partial z}(x, 0) = 0 \qquad \text{for } a < |x|\,, \tag{8.3}$$

and, as we replace the half-space by a large cylinder S_1,

$$\frac{\partial f}{\partial n} = 0 \qquad \text{on } S_1\,. \tag{8.4}$$

After the Kelvin inversion, we obtain

$$\nabla^2 g = 0 \tag{8.5}$$

in the indented circle C_3, and as boundary conditions

$$\frac{\partial g}{\partial n} = 0 \tag{8.6}$$

on the upper half-circle of C_3 and on the indentation C_4,

$$(1 - \cos\theta)\frac{\partial g}{\partial n} = 2\lambda g \tag{8.7}$$

on the lower half- circle of C_3, denoted again by S_l.

Eqs. (8.5) to (8.7) are the Euler-Lagrange equation and the natural boundary conditions of the variational problem, which consists of making the functional

$$L[g] = \iint |\nabla g|^2\, dA \Big/ \int_{S_l} \frac{2}{(1 - \cos\theta)}\, g^2\, dS$$

stationary. The double integral is extended over the indented circle (see Fig. 1).

The obvious choice for the coordinate functions for the Rayleigh-

Ritz method is now

$$f_n = \left(\frac{2r}{a}\right)^n \sin n\theta \,, \qquad n = 1, 2, \cdots$$

for the odd modes,

$$f_n = \left(\frac{2r}{a}\right)^n \cos n\theta \,, \qquad n = 0, 1, 2, \cdots$$

for the even modes.

There is no need to introduce a special function corresponding to eq. (5.2), since for the inversion by reciprocal radii in two dimensions a constant function transforms into a constant, and this term is included by setting $n = 0$ for the even modes. The limit $R_1 \to \infty$ also causes no problem.

A simple calculation furnishes for the matrix elements in

$$\det (A - \lambda B) = 0$$

the results

$$a_{jk} = j\pi \, \delta_{jk} \tag{8.8}$$

and

$$b_{jk} = (-1)^{j+k} a(C_{|j-k|} \pm C_{j+k})$$

with the upper sign for the even, the lower sign for the odd modes. The integrals

$$C_n = \int_0^{\pi/2} \frac{\cos n\varphi}{1 + \cos \varphi} \, d\varphi = \int_0^{\pi/4} \frac{\cos 2n\varphi}{\cos^2 \varphi} \, d\varphi$$

which appear in b_{jk} are readily evaluated as finite sums.

9. Numerical results. The numerical results for the upper bounds of the sloshing eigenvalues in a half-space are summarized in Table 1. They are based on the solution of a 16 by 16 matrix eigenvalue problem. Since the matrix A is diagonal in both the plane and the axi-symmetric case (cf. eqs. (6.3) and (8.8)), the eigenvalue problem is of the simplest type.

By observing the trend of the results as the order of the matrices A and B is increased to 16 by 16, and by applying an Aitken extrapolation, the eigenvalues can be improved, namely by a fraction of a percent for λ_1 and by a few percent for λ_5. However, the extrapolated values are no longer guaranteed to be upper bounds.

In Table 2 the present results are compared with previous results for three representative examples.

TABLE 1

	Planar case		axi-symmetric case						
	odd modes	even modes	$m = 0$	$m = 1$	$m = 2$	$m = 3$	$m = 4$	$m = 5$	$m = 6$
$a\lambda_1$	2.009	3.462	4.133	2.759	4.130	5.415	6.651	7.857	9.041
$a\lambda_2$	5.148	6.665	7.385	5.915	7.376	8.764	10.103	11.409	12.688
$a\lambda_3$	8.332	9.885	10.630	9.097	10.602	12.047	13.45	14.82	16.16
$a\lambda_4$	11.583	13.17	13.94	12.32	13.86	15.35	16.80	18.22	19.62
$a\lambda_5$	14.98	16.62	17.40	15.64	17.22	18.76	20.26	21.73	23.18

Upper bounds for the sloshing eigenvalues in a half-space, with strip-like aperture of width $2a$ in the planar case, and with circular aperture of radius a in the axi-symmetric case.

TABLE 2

	Present results	Ref. [2]	Ref. [4]
Planar case			
odd mode, $a\lambda_1$	2.009	2.018	2.006
axi-symmetric case			
$m = 1$, $a\lambda_3$	9.097	9.25	9.033
$m = 5$, $a\lambda_5$	21.73	—	21.14

Comparison of three typical examples with the upper bounds from [2] and [4].

Added in proof. The reader should also consult the paper by J. W. Miles, *On the eigenvalue problem for fluid sloshing in a half-space*, Z. Angew. Math. Phys., 23 (1972), 861-869.

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